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PH108 (Division 3) Lectures on TUESDAY & FRIDAY 1400-1525
SLOT NO 10A + 10B

\section*{Instructor (D3): Kantimay Das Gupta : kdasgupta@phy}

Reference texts:
D J Griffiths : Introduction to electromagnetism Feynman Lectures: vol 2

Vector Analysis (Schaum series) M Spiegel Mathematical methods : Pipes \& Harvil

Several other classic texts:
Panofsky and Philips
J D Jackson

ATTENDANCE : 80\% REQUIRED
EVALUATION Quiz1=15: Midsem=35: Quiz2=10 : Endsem=40

\section*{Electromagnetism}

Basic principles known for about 150 years.
Mature subject with a well defined structure.
Regime of validity well undesrtood.
Great success: explaining propagation \& generation of electromagnetic radiation, Forces of adhesion and cohesion.

First example of a classical field theory....particles and fields both carry energy and Momentum

Fails when we go to atomic scale
Gravity and electromagnetism are markedly different too, though both have " inverse square force" laws.

Two key questions:

Why do we use vectors?
Why do we use many co-ordinate systems?
Symmetry of the problem and the shape of the objects involved must be taken into account.

\section*{What is a field? What are the typical questions one asks?}

A quantity defined or measured over a certain area/volume of space.
Scalar field Temperature defined over a region \(T(x, y, z)\)
Vector field Electric, Magnetic field: \(\mathbf{E}(x, y, z) \quad \mathbf{B}(x, y, z)\) velocity of water \(\mathbf{v}(x, y, z)\) in a pipe, river, ocean

Matrix/Tensor field Stress, Strain inside a material like a concrete beam. With every point a matrix like object is associated.

A field is also like an object with a large number of degrees of freedom.

How is the field created? What is the "source"?
How does the field affect particles in it (Interaction of field with matter)?

\section*{A systematic way of handling co-ordinate systems : Part 1}

Many types of co-ordinates are needed, so that we can use the natural symmetry of a problem.

Equations would have the simplest form and minimum number of free variables if the co-ordinate system is chosen intelligently.

How to define a co-ordinate system?
Few typical systems:
Plane Polar
Spherical Polar
Cylindrical
Then how to define your own if you need?

Plane Polar \((r, \theta)\) in detail

STEP 1: Write down the relation with ( \(\mathrm{x}, \mathrm{y}\) ) co-ordinates
\[
\begin{aligned}
& x=r \cos \theta \\
& y=r \sin \theta
\end{aligned}
\]

STEP 2: Draw the co-ordinate grid


How do \(\mathrm{r}=\) constant lines look? How do \(\theta=\) constant lines look?

Plane Polar \((r, \theta)\) in detail
STEP 3: What happens when the independent variables are changed infinitesimally
\[
\begin{aligned}
& \delta x=\cos \theta \delta r-r \sin \theta \delta \theta \\
& \delta y=\sin \theta \delta r+r \cos \theta \delta \theta
\end{aligned}
\]

STEP 4: Which direction would we move, if only one variable was changed?
\[
\begin{aligned}
& \underline{\delta \theta=0} \\
& \boldsymbol{i} \delta x+\boldsymbol{j} \delta y=(\boldsymbol{i} \cos \theta+\boldsymbol{j} \sin \theta) \delta r \\
&=\boldsymbol{\epsilon}_{r} \delta r \\
& \begin{aligned}
\delta r=0
\end{aligned} \\
& \boldsymbol{i} \delta x+\boldsymbol{j} \delta y=(-\boldsymbol{i} \sin \theta+\boldsymbol{j} \cos \theta) r \delta \theta \\
&=\boldsymbol{\epsilon}_{\theta} r \delta \theta \\
& \delta \boldsymbol{r}=\boldsymbol{\epsilon}_{\boldsymbol{r}} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta
\end{aligned}
\]

\[
\boldsymbol{\epsilon}_{r} \cdot \boldsymbol{\epsilon}_{\theta}=0
\]

Curvilinear but still orthogonal

\section*{Plane Polar \((r, \theta)\) in detail}

STEP 5: What happens to the element of area?
i.e take a small step in the \(\epsilon_{r}\) direction and a small step in the \(\epsilon_{\theta}\) direction What is the "infinitesimal" area enclosed by these two perpendicular vectors?
\[
\begin{aligned}
d A & =\left|\boldsymbol{\epsilon}_{r} \times \boldsymbol{\epsilon}_{\theta}\right| \\
& =\left|\begin{array}{rc}
\cos \theta & \sin \theta \\
-r \sin \theta & r \cos \theta
\end{array}\right| \delta \theta \delta r \\
& =r \delta \theta \delta r
\end{aligned}
\]

STEP 6: What happens to the element of distance or arclength?
\[
\begin{aligned}
d s^{2} & =\delta \boldsymbol{r} . \delta \boldsymbol{r} \\
& =d r^{2}+r^{2} d \theta^{2}
\end{aligned}
\]

In orthogonal co-ordinates there will be no cross terms in the arclength expression.

\section*{Plane Polar \((r, \theta)\) in detail}

STEP 7: Now suppose a SCALAR function of co-ordinates is defined (like Temperature over a region), \(\mathrm{T}(\mathrm{x}, \mathrm{y}, \mathrm{z})\)

We change our position by a small VECTOR \(\delta \boldsymbol{r}\), and ask \(d T=\) ?
We want a function such that :
\[
\begin{aligned}
\delta T & =\frac{\partial T}{\partial r} \delta r+\frac{\partial T}{\partial \theta} \delta \theta \\
& =\left[\begin{array}{ll}
\text { some } & f n
\end{array}\right] \cdot \delta \boldsymbol{r} \\
& =\left[\begin{array}{ll}
\text { some } & f n
\end{array}\right] \cdot\left(\boldsymbol{\epsilon}_{r} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta\right)
\end{aligned}
\]
\(\left[\begin{array}{ll}\text { some } & f n\end{array}\right]=\boldsymbol{\epsilon}_{r} \frac{\partial T}{\partial r}+\boldsymbol{\epsilon}_{\theta} \frac{1}{r} \frac{\partial T}{\partial \theta}\)
The combination is called gradient
\[
\nabla=\boldsymbol{\epsilon}_{r} \frac{\partial}{\partial r}+\boldsymbol{\epsilon}_{\theta} \frac{1}{r} \frac{\partial}{\partial \theta}
\]

Gradient is the generalisation of the derivative in 1 dimension

Prove:
grad \(T\) is perpendicular to surfaces of constant T

What form would grad T take in cartesian coordinates?

\section*{Plane Polar \((r, \theta)\) in detail}

STEP 8: What are velocity and acceleration components, when a particle's motion Is described using polar co-ordinates?
\[
\begin{aligned}
\boldsymbol{v} & =\frac{\delta \boldsymbol{r}}{\delta t} \\
& =\frac{\delta}{\delta t}\left(\boldsymbol{\epsilon}_{r} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta\right) \\
& =\boldsymbol{\epsilon}_{r} \frac{d r}{d t}+\boldsymbol{\epsilon}_{\theta} r \frac{d \theta}{d t} \\
\boldsymbol{a} & =\frac{d}{d t}\left(\boldsymbol{\epsilon}_{r} \frac{d r}{d t}+\boldsymbol{\epsilon}_{\theta} r \frac{d \theta}{d t}\right)
\end{aligned}
\]

Unlike cartesian unit vectors the unit vectors here are not constant and must be differentiated themselves.
\[
\begin{aligned}
& \dot{\boldsymbol{\epsilon}}_{r}=? \\
& \dot{\boldsymbol{\epsilon}}_{\theta}=?
\end{aligned}
\]
\[
\binom{\boldsymbol{\epsilon}_{r}}{\boldsymbol{\epsilon}_{\theta}}=\left(\begin{array}{cc}
\cos \theta & \sin \theta \\
-\sin \theta & \cos \theta
\end{array}\right)\binom{\boldsymbol{i}}{\boldsymbol{j}}
\]
hence
\[
\binom{\dot{\boldsymbol{\epsilon}_{r}}}{\dot{\boldsymbol{\epsilon}_{\theta}}}=\dot{\theta}\left(\begin{array}{lr}
-\sin \theta & \cos \theta \\
-\cos \theta & -\sin \theta
\end{array}\right)\binom{\boldsymbol{i}}{\boldsymbol{j}}
\]

\section*{Plane Polar \((r, \theta)\) in detail}

STEP 8: What are velocity and acceleration components, when a particle's motion Is described using polar co-ordinates?
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\begin{aligned}
\boldsymbol{v} & =\frac{\delta \boldsymbol{r}}{\delta t} \\
& =\frac{\delta}{\delta t}\left(\boldsymbol{\epsilon}_{r} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta\right) \\
& =\boldsymbol{\epsilon}_{r} \frac{d r}{d t}+\boldsymbol{\epsilon}_{\theta} r \frac{d \theta}{d t} \\
\boldsymbol{a} & =\frac{d}{d t}\left(\boldsymbol{\epsilon}_{r} \frac{d r}{d t}+\boldsymbol{\epsilon}_{\theta} r \frac{d \theta}{d t}\right)
\end{aligned}
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\binom{\boldsymbol{\epsilon}_{r}}{\boldsymbol{\epsilon}_{\theta}}=\left(\begin{array}{cc}
\cos \theta & \sin \theta \\
-\sin \theta & \cos \theta
\end{array}\right)\binom{\boldsymbol{i}}{\boldsymbol{j}}
\]
hence
\[
\binom{\dot{\boldsymbol{\epsilon}_{r}}}{\dot{\boldsymbol{\epsilon}_{\theta}}}=\dot{\theta}\left(\begin{array}{lr}
-\sin \theta & \cos \theta \\
-\cos \theta & -\sin \theta
\end{array}\right)\binom{\boldsymbol{i}}{\boldsymbol{j}}
\]

Plane Polar \((r, \theta)\) in detail
Using the last two results:
\[
\begin{aligned}
& \binom{\dot{\epsilon}_{r}}{\boldsymbol{\epsilon}_{\theta}}=\dot{\theta}\left(\begin{array}{cc}
0 & 1 \\
-1 & 0
\end{array}\right)\binom{\boldsymbol{\epsilon}_{r}}{\boldsymbol{\epsilon}_{\theta}} \quad\left\{\begin{array}{l}
\dot{\epsilon}_{r}=\dot{\theta} \boldsymbol{\epsilon}_{\theta} \\
\dot{\epsilon}_{\theta}=-\dot{\epsilon_{r}}
\end{array}\right. \\
& \begin{aligned}
\boldsymbol{a}=\frac{d \boldsymbol{v}}{d t} & =\frac{d}{d t}\left(\boldsymbol{\epsilon}_{r} \dot{r}+\boldsymbol{\epsilon}_{\theta} r \dot{\theta}\right) \\
& =\boldsymbol{\epsilon}_{\theta} \dot{\theta} \dot{r}+\boldsymbol{\epsilon}_{r} \dot{r}-\boldsymbol{\epsilon}_{r} \dot{\theta} r \dot{\theta}+\boldsymbol{\epsilon}_{\theta} \dot{r} \dot{\theta}+\boldsymbol{\epsilon}_{\theta} r \ddot{\theta} \\
& =\left(\ddot{r}-\dot{\theta}^{2} r\right) \boldsymbol{\epsilon}_{r}+(r \ddot{\theta}+2 \dot{r} \dot{\theta}) \boldsymbol{\epsilon}_{\theta}
\end{aligned}
\end{aligned}
\]

What are the physical meanings of the various terms in the result for accelaration?

If the force on the particle is "central", then which quantity is conserved?

You are given an arbitrary vector in cartesian ( \(\boldsymbol{i} F_{x}+\boldsymbol{j} F_{y}\) ).
How will you go over to ( \(\boldsymbol{\epsilon}_{r} F_{r}+\boldsymbol{\epsilon}_{\theta} F_{\theta}\) )?
What can you say about the matrix connecting the two sets and the inverse relation?

Spherical Polar \((r, \theta, \phi)\) : two obvious examples


Spherical Polar (r, \(\theta, \phi)\)
\[
\begin{aligned}
& \hat{x} \times \hat{y}=\hat{z} \\
& x=r \sin \theta \cos \phi \\
& y=r \sin \theta \sin \phi \\
& z=r \cos \theta \\
& \text { Polar angle } \\
& \text { Azimuthal angle } \\
& \left(\begin{array}{l}
\delta x \\
\delta y \\
\delta z
\end{array}\right)=\left(\begin{array}{llc}
\sin \theta \cos \phi & r \cos \theta \cos \phi & -r \sin \theta \sin \phi \\
\sin \theta \sin \phi & r \cos \theta \sin \phi & r \sin \theta \cos \phi \\
\cos \theta & -r \sin \theta & 0
\end{array}\right)\left(\begin{array}{l}
\delta r \\
\delta \theta \\
\delta \phi
\end{array}\right)
\end{aligned}
\]

Spherical Polar \((r, \theta, \phi)\) : unit vectors, volume element, arc length
\[
\begin{array}{rrrrr}
\boldsymbol{\epsilon}_{r} & = & \sin \theta \cos \phi \boldsymbol{i} & + & \sin \theta \sin \phi \boldsymbol{j} \\
\boldsymbol{\epsilon}_{\theta} & = & \cos \theta \cos \theta \boldsymbol{k} \\
\boldsymbol{\epsilon}_{\phi} & = & -\sin \phi \boldsymbol{i} & + & \cos \theta \sin \phi \boldsymbol{j} \\
\hline & -\sin \theta \boldsymbol{k} \\
& \cos \phi \boldsymbol{j} &
\end{array}
\]

\section*{Can you invert this set of equations? It is easy!}

\[
\begin{aligned}
\delta \boldsymbol{r} & =\boldsymbol{\epsilon}_{r} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta+\boldsymbol{\epsilon}_{\phi} r \sin \theta \delta \phi \\
d s^{2} & =d r^{2}+r^{2} d \theta^{2}+r^{2} \sin ^{2} \theta d \phi^{2} \\
d V & =\left|\boldsymbol{\epsilon}_{r} \cdot\left(\boldsymbol{\epsilon}_{\theta} r\right) \times\left(\boldsymbol{\epsilon}_{\phi} r \sin \theta\right)\right| d r d \theta d \phi \\
& =r^{2} \sin \theta d r d \theta d \phi
\end{aligned}
\]

Spherical Polar \((r, \theta, \phi)\) : The area element
\[
\begin{aligned}
& \text { If } r=\text { constant (surface of a sphere) } \delta r=0 \\
& \begin{aligned}
d A & =\left|r \boldsymbol{\epsilon}_{\theta} \times r \sin \theta \boldsymbol{\epsilon}_{\phi}\right| d \theta d \phi \\
& =r^{2} \sin \theta d \theta d \phi
\end{aligned}
\end{aligned}
\]

If \(\theta=\) constant \(\delta \theta=0\)
\(d A=\left|\boldsymbol{\epsilon}_{r} \times r \sin \theta \boldsymbol{\epsilon}_{\boldsymbol{\phi}}\right| d r d \phi\)
\(=r \sin \theta d r d \phi\)

If \(\phi=\) constant (plane polar in a vertical plane) \(\delta \phi=0\)
\[
\begin{aligned}
d A & =\left|\boldsymbol{\epsilon}_{r} \times r \boldsymbol{\epsilon}_{\theta}\right| d r d \theta \\
& =r d r d \theta
\end{aligned}
\]

\section*{Q:}

Suppose you were confined on the surface of a sphere - but you were not told that. Would you be able to figure out?

Spherical Polar \((r, \theta, \phi)\) : velocity \& acceleration

We still need to express the derivatives \(\left(\dot{\boldsymbol{\epsilon}}_{r}, \dot{\boldsymbol{\epsilon}}_{\theta}, \dot{\boldsymbol{\epsilon}}_{\boldsymbol{\phi}}\right)\) in terms of \(\left(\boldsymbol{\epsilon}_{r}, \boldsymbol{\epsilon}_{\theta}, \boldsymbol{\epsilon}_{\phi}\right)\)
\[
\left(\begin{array}{l}
\dot{\boldsymbol{\epsilon}}_{r} \\
\dot{\boldsymbol{\epsilon}}_{\theta} \\
\dot{\boldsymbol{\epsilon}}_{\boldsymbol{\phi}}
\end{array}\right)=\dot{\boldsymbol{M}} \boldsymbol{M}^{T}\left(\begin{array}{c}
\boldsymbol{\epsilon}_{r} \\
\boldsymbol{\epsilon}_{\theta} \\
\boldsymbol{\epsilon}_{\phi}
\end{array}\right)
\]
\(\boldsymbol{M}=\left(\begin{array}{clc}\sin \theta \cos \phi & \sin \theta \sin \phi & \cos \theta \\ \cos \theta \cos \phi & \cos \theta \sin \phi & -\sin \theta \\ -\sin \phi & \cos \phi & 0\end{array}\right) \quad \boldsymbol{M}^{T}=\left(\begin{array}{lll}\sin \theta \cos \phi & \cos \theta \cos \phi & -\sin \phi \\ \sin \theta \sin \phi & \cos \theta \sin \phi & \cos \phi \\ \cos \theta & -\sin \theta & 0\end{array}\right)\)
\(\dot{\boldsymbol{M}}=\left(\begin{array}{lll}\cos \theta \cos \phi \dot{\theta}-\sin \theta \sin \phi \dot{\phi} & \cos \theta \sin \phi \dot{\theta}+\sin \theta \cos \phi \dot{\phi} & -\sin \theta \dot{\theta} \\ -\sin \theta \cos \phi \dot{\theta}-\cos \theta \sin \phi \dot{\phi} & -\sin \theta \sin \phi \dot{\theta}+\cos \theta \cos \phi \dot{\phi} & -\cos \theta \dot{\theta} \\ -\cos \phi \dot{\phi} & -\sin \phi \dot{\phi} & 0\end{array}\right)\)

Spherical Polar \((r, \theta, \phi)\) : velocity \& acceleration
This appears very messy! But if you work through the matrix multiplication then:
\[
\begin{aligned}
\left(\begin{array}{c}
\dot{\boldsymbol{\epsilon}_{r}} \\
\dot{\boldsymbol{\epsilon}_{\theta}} \\
\dot{\boldsymbol{\epsilon}_{\phi}}
\end{array}\right) & =\dot{\boldsymbol{M}} \boldsymbol{M}^{\boldsymbol{T}}\left(\begin{array}{c}
\boldsymbol{\epsilon}_{r} \\
\boldsymbol{\epsilon}_{\theta} \\
\boldsymbol{\epsilon}_{\phi}
\end{array}\right) \\
& =\left(\begin{array}{ccc}
0 & \dot{\theta} & \sin \theta \dot{\phi} \\
-\dot{\theta} & 0 & \cos \theta \dot{\phi} \\
-\sin \theta \dot{\phi} & -\cos \theta \dot{\phi} & 0
\end{array}\right)\left(\begin{array}{l}
\boldsymbol{\epsilon}_{r} \\
\boldsymbol{\epsilon}_{\theta} \\
\boldsymbol{\epsilon}_{\phi}
\end{array}\right)
\end{aligned}
\]

The result is remarkably simple.
Why are the diagonal terms zero? Can you see the physical implication?
Notice that the matrix connecting the two vectors is anti-symmetric.
This was also the case in the plane polar co-ordinates. But we didn't mention it there.

The problem for velocity and acceleration components can now be completed...

Spherical Polar \((r, \theta, \phi)\) : The gradient
If we have a function \(T(r, \theta, \phi)\) then we want
\[
\begin{aligned}
\delta T & =\frac{\partial T}{\partial r} \delta r+\frac{\partial T}{\partial \theta} \delta \theta+\frac{\partial T}{\partial \phi} \delta \phi \\
& =\nabla T . \delta \boldsymbol{r}
\end{aligned}
\]
since
\(\delta \boldsymbol{r}=\boldsymbol{\epsilon}_{\boldsymbol{r}} \delta r+\boldsymbol{\epsilon}_{\theta} r \delta \theta+\boldsymbol{\epsilon}_{\boldsymbol{\phi}} r \sin \theta \delta \phi\)
we must have
\(\nabla T=\boldsymbol{\epsilon}_{r} \frac{\partial T}{\partial r}+\boldsymbol{\epsilon}_{\theta} \frac{1}{r} \frac{\partial T}{\partial \theta}+\boldsymbol{\epsilon}_{\phi} \frac{1}{r \sin \theta} \frac{\partial T}{\partial \phi}\)

Spherical Polar \((r, \theta, \phi)\) : velocity \& acceleration
You should be able to show the following now:
\[
\begin{aligned}
\boldsymbol{v}= & \boldsymbol{\epsilon}_{\boldsymbol{r}} \dot{r}+\boldsymbol{\epsilon}_{\boldsymbol{\theta}} r \dot{\theta}+\boldsymbol{\epsilon}_{\phi} r \sin \theta \dot{\phi} \\
\boldsymbol{a}= & \boldsymbol{\epsilon}_{\boldsymbol{r}}\left(\ddot{r}-r \dot{\theta}^{2}-r \dot{\phi}^{2} \sin ^{2} \theta\right)+ \\
& \boldsymbol{\epsilon}_{\boldsymbol{\theta}}\left(r \ddot{\theta}+2 \dot{r} \dot{\theta}-r \dot{\phi}^{2} \sin \theta \cos \theta\right)+ \\
& \boldsymbol{\epsilon}_{\boldsymbol{\phi}}(r \ddot{\phi} \sin \theta+2 \dot{r} \dot{\phi} \sin \theta+2 r \dot{\theta} \dot{\phi} \cos \theta)
\end{aligned}
\]

We now have all the necessary bits to solve dynamical problems in this co-ordinate

Cylindrical polar ( \(\rho, \theta, z\) ) : unit vectors

\[
\begin{aligned}
& x=\rho \cos \phi \\
& y=\rho \sin \phi \\
& z=z
\end{aligned}
\]

Wires, co-axial cables, Pipes etc.

Cylindrical polar ( \(\rho, \theta, z\) ) : length, area and volume elements
\(d s^{2}=d \rho^{2}+\rho^{2} d \phi^{2}+d z^{2}\)
\(d A=\rho d \rho d \phi \quad z=\) constant
\(d A=\rho d \phi d z \quad \rho=\) constant
\(d A=d \rho d z \quad \phi=\) constant
volume
\(d V=\rho d \rho d \phi d z\)
gradient
\(\nabla=\boldsymbol{\epsilon}_{\rho} \frac{\partial}{\partial \rho}+\boldsymbol{\epsilon}_{\phi} \frac{1}{\rho} \frac{\partial}{\partial \phi}+\boldsymbol{\epsilon}_{z} \frac{\partial}{\partial z}\)

Follow exactly the same process as we did for spherical polar...

Writing the basic information about orthogonal co-ordinates.... \(d \boldsymbol{r}=\boldsymbol{\epsilon}_{\mathbf{1}} h_{1} d u_{1}+\boldsymbol{\epsilon}_{\mathbf{2}} h_{2} d u_{2}+\boldsymbol{\epsilon}_{\mathbf{3}} h_{3} d u_{3}\)
\(d s^{2}=\) ?
\(d V=\) ?

A shorthand compact way of writing co-ordinates
\(d \boldsymbol{r}=\sum \boldsymbol{\epsilon}_{\boldsymbol{i}} h_{i} d u_{i}\)

Summation convention : REPEATED INDEX IMPLIES SUMMATION
\(d \boldsymbol{r}=\boldsymbol{\epsilon}_{i} h_{i} d u_{i}\)

Flux and circulation


The volume of water flowing out through the SURFACE per unit time
\[
\oiint \boldsymbol{v} \cdot d \boldsymbol{S}
\]

The shape of the surface can be arbitrary

\section*{dS points OUTWARD}

This has a unique meaning only if the surface is closed.

\section*{Flux and circulation}

Consider a vector \(\boldsymbol{F}\)
Is it possible to have a function \(\boldsymbol{X}(\boldsymbol{F})\) such that \(X(\boldsymbol{F}) d V=\boldsymbol{F} . d \boldsymbol{S}\)

Flux through BACK

Top


BOTTOM
\[
f_{B}=-F_{1} h_{2} \delta u_{2} h_{3} \delta u_{3}
\]

Flux through FRONT
\[
\begin{aligned}
f_{F} & =F_{1} h_{2} \delta u_{2} h_{3} \delta u_{3} \\
& +\frac{\partial}{\partial u_{1}}\left(F_{1} h_{2} \delta u_{2} h_{3} \delta u_{3}\right) \delta u_{1}
\end{aligned}
\]
\[
f_{B}+f_{F}=\left[\frac{\partial}{\partial u_{1}}\left(F_{1} h_{2} h_{3}\right)\right] \delta u_{1} \delta u_{2} \delta u_{3}
\]
!! BE VERY CLEAR ABOUT THE SIGN OF EACH QUANTITY !!

Flux and circulation

The LEFT + RIGHT pair gives
\(f_{L}+f_{R}=\left[\frac{\partial}{\partial u_{2}}\left(F_{2} h_{1} h_{3}\right)\right] \delta u_{1} \delta u_{2} \delta u_{3}\)
The BOTTOM + TOP pair gives
\(f_{\text {Botom }}+f_{\text {Top }}=\left[\frac{\partial}{\partial u_{3}}\left(F_{3} h_{1} h_{2}\right)\right] \delta u_{1} \delta u_{2} \delta u_{3}\)
\(f_{\text {TOTAL }}=\left[\frac{\partial}{\partial u_{1}}\left(F_{1} h_{2} h_{3}\right)+\frac{\partial}{\partial u_{2}}\left(F_{2} h_{1} h_{3}\right)+\frac{\partial}{\partial u_{3}}\left(F_{3} h_{1} h_{2}\right)\right] \delta u_{1} \delta u_{2} \delta u_{3}\)
\(\frac{\boldsymbol{F} \cdot \delta \boldsymbol{S}}{\delta V}=\frac{1}{h_{1} h_{2} h_{3}}\left[\frac{\partial}{\partial u_{1}}\left(F_{1} h_{2} h_{3}\right)+\frac{\partial}{\partial u_{2}}\left(F_{2} h_{3} h_{1}\right)+\frac{\partial}{\partial u_{3}}\left(F_{3} h_{1} h_{2}\right)\right]\)
Now break a finite volume into small volume elements
Flux from neighbouring walls of two infinitesimal volume elements will cancel
Only faces which form the part of the boundary of the volume will not cancel

This function is called DIVERGENCE, denoted by \(\boldsymbol{\nabla} . \boldsymbol{F}\) \(\oiiint \boldsymbol{\nabla} \cdot \boldsymbol{F} d V=\oiint \boldsymbol{F} . d \boldsymbol{S}\)
Called Gauss 's theorem
Divergence of a vector is a scalar quantity
In Cartesian:
\(\nabla . \boldsymbol{F}=\boldsymbol{i} \frac{\partial F_{x}}{\partial x}+\boldsymbol{i} \frac{\partial F_{y}}{\partial y}+\boldsymbol{k} \frac{\partial F_{z}}{\partial z}\)
"divergence" should convey a visual picture of the Vector field.... What is it?

How should a vector field look around points of stable/unstable equilibrium?

In Spherical polar:
\(\nabla \cdot \boldsymbol{F}=\frac{1}{r^{2} \sin \theta}\left[\frac{\partial}{\partial r}\left(r^{2} \sin \theta F_{r}\right)+\frac{\partial}{\partial \theta}\left(r \sin \theta F_{\theta}\right)+\frac{\partial}{\partial \phi}\left(r F_{\phi}\right)\right]\)
In cylindrical polar
\(\nabla \cdot \boldsymbol{F}=\frac{1}{\rho}\left[\frac{\partial}{\partial \rho}\left(\rho F_{\rho}\right)+\frac{\partial}{\partial \phi}\left(F_{\phi}\right)+\frac{\partial}{\partial z}\left(\rho F_{z}\right)\right]\)
Divergence and continuity equation....

Flux and circulation

\(\oiint \boldsymbol{F} . d \boldsymbol{S}\) identifies a distinctive field pattern.

Another possible one is a circulating pattern.


When will \(\oint \boldsymbol{F} . d \boldsymbol{l}\) be nonzero ?

Flux and circulation
Consider two arbitray infinitesimal displacements

\[
\begin{aligned}
& \delta \boldsymbol{r}^{\alpha}=\boldsymbol{\epsilon}_{\mathbf{1}} h_{1} \delta u_{1}^{\alpha}+\boldsymbol{\epsilon}_{\mathbf{2}} h_{2} \delta u_{2}^{\alpha}+\boldsymbol{\epsilon}_{3} h_{3} \delta u_{3}^{\alpha} \\
& \delta \boldsymbol{r}^{\beta}=\boldsymbol{\epsilon}_{\mathbf{1}} h_{1} \delta u_{1}^{\beta}+\boldsymbol{\epsilon}_{\mathbf{2}} h_{2} \delta u_{2}^{\beta}+\boldsymbol{\epsilon}_{\mathbf{3}} h_{3} \delta u_{3}^{\beta}
\end{aligned}
\]

The vector field is \(\boldsymbol{F}\).
Is it possible to have a function \(\boldsymbol{X}(\boldsymbol{F})\) such that
\[
X(\boldsymbol{F}) . \delta \boldsymbol{S}=\sum_{\substack{\text { peri- } \\ \text { meter }}} \boldsymbol{F} . \delta \boldsymbol{l}
\]

If possible then this function will connect some characteristics of inside points with the boundary

Flux and circulation
\[
\begin{aligned}
d \boldsymbol{S}=\delta \boldsymbol{r}^{\alpha} \times \delta \boldsymbol{r}^{\beta}= & \left|\begin{array}{lll}
\boldsymbol{\epsilon}_{1} & \boldsymbol{\epsilon}_{2} & \boldsymbol{\epsilon}_{3} \\
h_{1} \delta u_{1}^{\alpha} & h_{2} \delta u_{2}^{\alpha} & h_{3} \delta u_{3}^{\alpha} \\
h_{1} \delta u_{1}^{\beta} & h_{2} \delta u_{2}^{\beta} & h_{3} \delta u_{3}^{\beta}
\end{array}\right| \\
X(\boldsymbol{F}) . d \boldsymbol{S}= & X_{1} h_{2} h_{3}\left[\delta u_{2}^{\alpha} \delta u_{3}^{\beta}-\delta u_{3}^{\alpha} \delta u_{2}^{\beta}\right] \\
& -X_{2} h_{1} h_{3}\left[\delta u_{1}^{\alpha} \delta u_{3}^{\beta}-\delta u_{3}^{\alpha} \delta u_{1}^{\beta}\right] \\
& +X_{3} h_{1} h_{2}\left[\delta u_{1}^{\alpha} \delta u_{2}^{\beta}-\delta u_{2}^{\alpha} \delta u_{1}^{\beta}\right]
\end{aligned}
\]

Try writing RHS in this form and compare.
The co-efficients of the arbitrary displacments must agree
!! BE VERY CLEAR ABOUT THE SIGN OF EACH QUANTITY !!

Flux and circulation
Consider the pair of paths \((1 \rightarrow 2)\) and \((3 \rightarrow 4)\)
\(\boldsymbol{F} .\left.\delta \boldsymbol{l}\right|_{1 \rightarrow 2}=F_{1} h_{1} \delta u_{1}^{\alpha}+F_{2} h_{2} \delta u_{2}^{\alpha}+F_{3} h_{3} \delta u_{3}^{\alpha}\)
\(\boldsymbol{F} .\left.\delta \boldsymbol{l}\right|_{3 \rightarrow 4}=\left[F_{i} h_{i}+\left(\boldsymbol{\nabla} F_{i} h_{i}\right) \cdot \delta \boldsymbol{r}^{\beta}\right]\left(-\delta u_{i}^{\alpha}\right) \quad(i=1,2,3)\)

Write contributions from \(\boldsymbol{F} .\left.\delta \boldsymbol{l}\right|_{2 \rightarrow 3} \& \boldsymbol{F} .\left.\delta \boldsymbol{l}\right|_{4 \rightarrow 1}\) similarly.
Full path gives: \(\quad\left(\boldsymbol{\nabla} \boldsymbol{F} . \delta r^{\beta}\right) . \delta r^{\alpha}-\left(\boldsymbol{\nabla} \boldsymbol{F} . \delta r^{\alpha}\right) \cdot \delta r^{\beta}\)

!! BE VERY CLEAR ABOUT THE SIGN OF EACH QUANTITY !!

Flux and circulation
Now compare the co-efficient of \(\delta u_{2}^{\alpha} \delta u_{3}^{\beta}-\delta u_{3}^{\alpha} \delta u_{2}^{\beta}\)
We need to put \(i=3, k=2\) and then \(i=2, k=3\)
this gives \(\quad X_{1} h_{2} h_{3}=\left[\frac{\partial F_{3} h_{3}}{\partial u_{2}}-\frac{\partial F_{2} h_{2}}{\partial u_{3}}\right]\)
So \(\quad X(\boldsymbol{F})=\frac{1}{h_{1} h_{2} h_{3}}\left|\begin{array}{lll}h_{1} \boldsymbol{\epsilon}_{1} & h_{2} \boldsymbol{\epsilon}_{2} & h_{3} \boldsymbol{\epsilon}_{3} \\ \frac{\partial}{\partial u_{1}} & \frac{\partial}{\partial u_{2}} & \frac{\partial}{\partial u_{3}} \\ h_{1} F_{1} & h_{2} F_{2} & h_{3} F_{3}\end{array}\right| \equiv\left\{\begin{array}{l}\boldsymbol{\nabla} \times \boldsymbol{F} \\ \operatorname{curl} \boldsymbol{F} \\ \operatorname{rot} \boldsymbol{F}\end{array}\right.\)
We have \(\iint \nabla \times \boldsymbol{F} . d \boldsymbol{S}=\oint \boldsymbol{F} . d \boldsymbol{l} \quad\) (called Stoke's theorem)
Now break a finite surface into small area elements

Line integral from neighbouring perimeters of two infinitesimal area elements will cancel

Only line segments which form the part of the perimeter will not cancel

Flux and circulation : Which surface?


Any surface with the same bounding edge will work.
Curl F over any closed surface should be zero. WHY?
Divergence of a curl = ?
Curl of a gradient \(=\) ?

\section*{Multiple vector products : \(\varepsilon-\delta\) notation (Levi-Civita)}

Write the dot product as \(\boldsymbol{A} \cdot \boldsymbol{B}=\delta_{i j} A_{i} B_{j} \quad\) where \(\delta_{i j}=\left\{\begin{array}{lll}1 & \text { if } i=j \\ 0 & \text { if } i \neq j\end{array}\right.\)
Write the cross product as
\(\boldsymbol{A} \times\left.\boldsymbol{B}\right|_{i}=\epsilon_{i j k} A_{j} B_{k} \quad\) where \(\epsilon_{i j k}=\left\{\begin{aligned} 1 & \text { if ijk is an even permutation of } 123 \\ -1 & \text { if ijk is an odd permutation of } 123 \\ 0 & \text { otherwise }\end{aligned}\right.\)
Convince yourself that \(\epsilon_{i j k}=\boldsymbol{\epsilon}_{\boldsymbol{i}} . \boldsymbol{\epsilon}_{\boldsymbol{j}} \times \boldsymbol{\epsilon}_{\boldsymbol{k}}\)
This works with operators also: with \(x_{i}\) for \(x, y, z\)
\(\nabla . \boldsymbol{A}=\frac{\partial A_{i}}{\partial x_{i}}\)
\(\nabla \times\left.\boldsymbol{A}\right|_{i}=\epsilon_{i j k} \frac{\partial A_{j}}{\partial x_{k}}\)

Notice how the summation convention on repeated indices have been used.

Q: How does it help?

\section*{Multiple vector products : \(\varepsilon-\delta\) notation (Levi-Civita)}

Consider a vector triple product \(\quad \boldsymbol{A} \times(\boldsymbol{B} \times \boldsymbol{C})\)
\[
\begin{aligned}
\boldsymbol{A} \times\left.(\boldsymbol{B} \times \boldsymbol{C})\right|_{i} & =\epsilon_{i j k} A_{j}(\boldsymbol{B} \times \boldsymbol{C})_{k} \\
& =\epsilon_{i j k} A_{j}\left(\epsilon_{k p q} B_{p} C_{q}\right) \\
& =\epsilon_{k i j} \epsilon_{k p q} A_{j} B_{p} C_{q}
\end{aligned}
\]

What to do with a product like \(\epsilon_{i j k} \epsilon_{k p q}\) ?
This is either -1 or 0 or 1
kpq itself is some permutation of the sequence Imn, otherwise the WITHOUT summation, \(\begin{aligned} & \text { the equuence } \mathrm{mn} \text {, } \\ & \text { product will vanish. }\end{aligned}\)
we can write for generic product term: (why?)
\(\epsilon_{l m n} \epsilon_{k p q}=\begin{array}{r}\delta_{l k} \delta_{m p} \delta_{n q}+\delta_{l p} \delta_{m q} \delta_{n k}+\delta_{l q} \delta_{m k} \delta_{n p} \\ -\delta_{l k} \delta_{m q} \delta_{n p}-\delta_{l p} \delta_{m k} \delta_{n q}-\delta_{l q} \delta_{m p} \delta_{n k}\end{array}\)
odd/even
permutations
\begin{tabular}{ll}
\(k p q\) & + \\
\(k q p\) & - \\
\(p k q\) & - \\
\(p q k\) & + \\
\(q k p\) & + \\
\(q p k\) & -
\end{tabular}

\section*{Multiple vector products : \(\varepsilon-\delta\) notation (Levi-Civita)}
\[
\begin{aligned}
\epsilon_{k i j} \epsilon_{k p q}= & \delta_{k k} \delta_{i p} \delta_{j q}+\delta_{k p} \delta_{i q} \delta_{j k}+\delta_{k q} \delta_{i k} \delta_{j p} \quad \text { sum over k } \\
& -\delta_{k k} \delta_{i q} \delta_{j p}-\delta_{k p} \delta_{i k} \delta_{j q}-\delta_{k q} \delta_{i p} \delta_{j k} \\
= & \delta_{k k}\left(\delta_{i p} \delta_{j q}-\delta_{i q} \delta_{j p}\right)+\delta_{k p}\left(\delta_{i q} \delta_{j k}-\delta_{i k} \delta_{j q}\right)+\delta_{k q}\left(\delta_{i k} \delta_{j p}-\delta_{i p} \delta_{j k}\right) \\
= & 3\left(\delta_{i p} \delta_{j q}-\delta_{i q} \delta_{j p}\right)+\left(\delta_{i q} \delta_{j p}-\delta_{i p} \delta_{j q}\right)+\left(\delta_{i q} \delta_{j p}-\delta_{i p} \delta_{j q}\right) \\
= & \delta_{i p} \delta_{j q}-\delta_{i q} \delta_{j p}
\end{aligned}
\]

Using the last result ( with \(i=p\) )
\(\epsilon_{k i j} \epsilon_{k i q}=\delta_{i i} \delta_{j q}-\delta_{i q} \delta_{j i}\) sum over k and i
\[
=3 \delta_{j q}-\delta_{j q}
\]
\[
=2 \delta_{j q}
\]

Successive summation over indices.

The first sum is most frequently encountered. It allows you to write a cross product in terms of dot product like terms

Using the last result ( with \(j=q\) )
\[
\begin{aligned}
\epsilon_{k i j} \epsilon_{k j} & =2 \delta_{j j} \quad \text { sum over k, and } j \\
& =6
\end{aligned}
\]

\section*{Multiple vector products : \(\varepsilon-\delta\) notation (Levi-Civita)}

\section*{TRIPLE PRODUCTS}
(1) \(\mathbf{A} \cdot(\mathbf{B} \times \mathbf{C})=\mathbf{B} \cdot(\mathbf{C} \times \mathbf{A})=\mathbf{C} \cdot(\mathbf{A} \times \mathbf{B})\)
(2) \(\mathbf{A} \times(\mathbf{B} \times \mathbf{C})=\mathbf{B}(\mathbf{A} \cdot \mathbf{C})-\mathbf{C}(\mathbf{A} \cdot \mathbf{B})\)

Should be able to prove all of these easily....
(list from the last page of Griffith's book)

\section*{PRODUCT RULES}
(3) \(\quad \nabla(f g)=f(\nabla g)+g(\nabla f)\)
(4) \(\boldsymbol{\nabla}(\mathbf{A} \cdot \mathbf{B})=\mathbf{A} \times(\boldsymbol{\nabla} \times \mathbf{B})+\mathbf{B} \times(\boldsymbol{\nabla} \times \mathbf{A})+(\mathbf{A} \cdot \nabla) \mathbf{B}+(\mathbf{B} \cdot \boldsymbol{\nabla}) \mathbf{A}\)
(5) \(\boldsymbol{\nabla} \cdot(f \mathbf{A})=f(\nabla \cdot \mathbf{A})+\mathbf{A} \cdot(\nabla f)\)
(6) \(\boldsymbol{\nabla} \cdot(\mathbf{A} \times \mathbf{B})=\mathbf{B} \cdot(\boldsymbol{\nabla} \times \mathbf{A})-\mathbf{A} \cdot(\boldsymbol{\nabla} \times \mathbf{B})\)
(7) \(\nabla \times(f \mathbf{A})=f(\nabla \times \mathbf{A})-\mathbf{A} \times(\nabla f)\)
(8) \(\boldsymbol{\nabla} \times(\mathbf{A} \times \mathbf{B})=(\mathbf{B} \cdot \nabla) \mathbf{A}-(\mathbf{A} \cdot \boldsymbol{\nabla}) \mathbf{B}+\mathbf{A}(\boldsymbol{\nabla} \cdot \mathbf{B})-\mathbf{B}(\boldsymbol{\nabla} \cdot \mathbf{A})\)

SECOND DERIVATIVES
.(9) \(\nabla \cdot(\nabla \times A)=0 \ldots\)
(10) \(\nabla \times(\nabla f)=0\)
(11) \(\boldsymbol{\nabla} \times(\boldsymbol{\nabla} \times \mathbf{A})=\nabla(\nabla \cdot \mathbf{A})-\nabla^{2} \mathbf{A}\)

\section*{Co-ordinate systems, transformations etc : few problems}

You are given the set of co-ordinate transformation equations, \(\operatorname{say}(r, \theta, \phi) \rightarrow(x, y, z)\) What is the value of the following determinant \(|J|\), where
\(J=\left\lvert\,\)\begin{tabular}{lll}
\(\frac{\partial x}{\partial r}\) & \(\frac{\partial y}{\partial r}\) & \(\frac{\partial z}{\partial r}\) \\
\(\frac{\partial x}{\partial \theta}\) & \(\frac{\partial y}{\partial \theta}\) & \(\frac{\partial z}{\partial \theta}\) \\
\(\frac{\partial x}{\partial \phi}\) & \(\frac{\partial y}{\partial \phi}\) & \(\frac{\partial z}{\partial \phi}\)
\end{tabular}\(\quad\)\begin{tabular}{l} 
What is the physical significance of the result? \\
If the transformation was non-orthogonal, \\
would the same significance still hold ? \\
Can you makeuse of \(J\), for calculating the inverse partial derivatives
\end{tabular}\right.

When is a vector field uniquely specified?

Suppose \(\quad \boldsymbol{A}=\boldsymbol{\nabla} \times \boldsymbol{F}\) is specified.
Can we solve for \(\boldsymbol{F}\) ?
\(\phi(x, y, z)\) is any scalar field
NO. \(\quad \boldsymbol{F}^{\prime}=\boldsymbol{F}+\nabla \phi \quad\) will give same result.
Suppose \(\psi=\boldsymbol{\nabla} . \boldsymbol{F}\) is specified.....
\(B(x, y, z)\) is any vector field
NO. \(\quad \boldsymbol{F}^{\prime}=\boldsymbol{F}+\boldsymbol{\nabla} \times \boldsymbol{B} \quad\) will give same result.
If \(\boldsymbol{\nabla} \times \boldsymbol{F}\) AND \(\boldsymbol{\nabla}\). \(\boldsymbol{F}\) aregiven.
Then \(\boldsymbol{F}\) canbesolved for,

HELMHOLTZ'S THEOREM
provided boundary conditions are also given.
Maxwell's equations do precisely this.

Gauss's law: Flux of Electric field through a closed surface


Gauss's law: Flux of Electric field through a closed surface


Gauss's law: Flux of Electric field through a closed surface
\(\int_{\text {surface }} \boldsymbol{E} . d \boldsymbol{S}=\int_{\text {vol }} \boldsymbol{\nabla} . \boldsymbol{E} d \tau=\int_{\text {vol }} \frac{\rho(\boldsymbol{r})}{\epsilon_{0}} d \tau \quad\)\begin{tabular}{l} 
Q: Why do we write \\
This rather than leave \\
Coulomb force law \\
as it is?
\end{tabular}
So \(\boldsymbol{\nabla} . \boldsymbol{E}=\frac{\rho(\boldsymbol{r})}{\epsilon_{0}}\)
Also \(\nabla^{2} V=-\frac{\rho(\boldsymbol{r})}{\epsilon_{0}}\) since \(\boldsymbol{E}=-\boldsymbol{\nabla} V\)\begin{tabular}{l} 
Only when there \\
is no time varying \\
magnetic field. \\
Same as saying \\
Curl \(=0\)
\end{tabular}

This form allows one to use the symmetry of a problem more easily (e.g. sphere, infinite sheet, wire etc.)

Is valid even when charges are in motion.
Q: What is the problem with moving charges and Coulomb's law?

Fun question: If the world was 2-dimensional what would Coulomb's law be like? (Don't take it too seriously!)

How "exact" is the inverse square force law?

The cancellation of the \(1 / \mathrm{r}^{2}\) came from two sources:
1. The geometrical growth of the area subtended by a small solid angle (geometry)
2. The nature of the coulomb force law (experimental observation)

If the force varied as \(1 / /^{2.0001}\), what observational consequence would it have?

Gauss's law: divergence of \(1 / r^{2}\) and the Dirac delta function
Do the following integration over a sphere
\(\int_{\text {vol }} \nabla \cdot \frac{\hat{\boldsymbol{r}}}{r^{2}} d \tau=\int_{\text {surface }} \frac{\hat{\boldsymbol{r}}}{r^{2}} \cdot d \boldsymbol{S}=\int_{0}^{2 \pi} d \phi \int_{0}^{\pi} \sin \theta d \theta=4 \pi\)
But calculating the divergence explicitly
\(\nabla \cdot \frac{\hat{r}}{r^{2}}=\frac{\partial}{\partial x}\left(\frac{x}{r^{3}}\right)+\frac{\partial}{\partial y}\left(\frac{y}{r^{3}}\right)+\frac{\partial}{\partial z}\left(\frac{z}{r^{3}}\right)=0\)

The inconsistency comes from the singularity at \(\mathrm{r}=0\)

Consider a simpler example: the step function
\[
x(t)= \begin{cases}0 & t<0 \\ 1 & t>0\end{cases}
\]

Integrable
singularity at \(\mathbf{r}=0\)
\(\frac{d x}{d t}=\) ? looks like zero everywhere but must satisfy
\[
\int_{0-|\epsilon|}^{0+|\epsilon|}\left(\frac{d x}{d t}\right) d t=x(|\mathrm{E}|)-x(-|\epsilon|)=1 \quad \forall \epsilon \neq 0
\]


Such integrable singularities are treated by defining the \(\delta\) function

Gauss's law: divergence of \(1 / r^{2}\) and the Dirac delta function

Such "functions" can only be defined by specifying their behaviour inside an integral. You cannot really plot such functions because they are inherently singular.
\(\int_{\mathrm{a}}^{\mathrm{b}} \delta\left(x-x_{0}\right) f(x) d x= \begin{cases}f\left(x_{0}\right), & \text { if } x_{0} \text { is within the limits } \\ 0 & \text { otherwise }\end{cases}\)

Visualise this as a huge spike at \(x=x_{0}\) only,
Gets higher but narrower keeping the area under it, same.
Picks out the value of any \(f(x)\) at the spike Several other ways to define \(\delta(x)\) as a limit For our purpose, we will need to use
\(\nabla \cdot \frac{\hat{\boldsymbol{r}}}{r^{2}}=4 \pi \delta(\boldsymbol{r})\)

\section*{Fourier \& Cauchy had introduced such "functions" Before.}

In physics texts It is generally associated with Dirac

Gauss's law: divergence of \(1 / r^{2}\) and the Dirac delta function
Someother integral representations of \(\delta\) function
\(\frac{1}{2 \pi} \int_{-\infty}^{\infty} d p e^{i p\left(x-x_{0}\right)}=\delta\left(x-x_{0}\right) \quad \begin{aligned} & \text { Frequently used in } \\ & \text { Quantum mechanics }\end{aligned}\)
\(\lim _{a \rightarrow 0} \frac{1}{a \sqrt{ } \pi} e^{-x^{2} / a^{2}}=\delta(x)\)

Try proving... (hint: change of variables)
\(\delta(-x)=\delta(x) \quad \alpha\) is any constant
\(\delta(\alpha x)=\frac{\delta(x)}{|\alpha|}\)

The wikipedia article is excellent! Read it.

\section*{Electric field in some simple situations (symmetry + Gauss's Law)}

No charge distribution is really infinite!
Very close to the surface/
Objects with very large aspect ratio
infinite
approximation
is useful

Spherically symmeteric charge distribution

\(E_{\theta}=0\) and \(E_{\phi}=0 . W h y ?\)
rotate about \(z\)-axis
The field at the tip of \(\boldsymbol{r}\), must be same at all points visited by \(\boldsymbol{r}\)
\(\oint \boldsymbol{E} . \boldsymbol{d} \boldsymbol{l}=0 \quad \rightarrow E_{\phi}=0\)
Rotate about \(x\)-axis : show \(E_{\theta}=0\)
\[
E_{r} 4 \pi R^{2}=\frac{1}{\epsilon_{0}} \int_{0}^{R} \rho(r) 4 \pi r^{2} d r
\]

Electric field in some simple situations（symmetry＋Gauss＇s Law）

Long narrow wire type charge distribution
```

$E_{\phi}=0$ and $E_{z}=0$. Why?
rotate about $z$-axis
くニニ- - コロニン $\quad$ The field at the tip of $\boldsymbol{P}$, must be same
at all points visited by $\boldsymbol{P}$
$\oint \boldsymbol{E} . \boldsymbol{d} \boldsymbol{l}=0 \quad \rightarrow E_{\phi}=0$
flipabout $z$-axis : show $E_{z}=0$
There is nothing to chose $z$ from $-z$

$$
E_{\rho} 2 \pi \rho=\frac{1}{\epsilon_{0}} \lambda \quad \lambda: \text { charge per unit length }
$$

```

\section*{Electric field in some simple situations (symmetry + Gauss's Law)}

\section*{Infinite sheet of charge}


Give a symmetry argument to show that \(E(z)=-E(-z)\) must hold.
Flip the "sheet" switching "topside" and "bottom-side". What should happen?

Why cannot there be a \(E_{\|}\)component?
Rotate the sheet about any point Translate sheet by any in plane vector Field cannot change.
Not possible if there is finite component.
\(E_{z}=\left\{\begin{aligned} \frac{\sigma}{2 \epsilon_{0}} & : z>0 \\ -\frac{\sigma}{2 \epsilon_{0}} & : z<0\end{aligned}\right.\)
Easy to extend to a sheet with finite thickness....work it out.

Superpose the field of two parallel plates, calculate capacitance

Some extra charge is placed in a conductor : why does it go to the surface?
\[
\begin{aligned}
& \begin{array}{rlr}
\nabla . \boldsymbol{E} & =\frac{\rho}{\epsilon_{0}} & \text { Ohm's Law } \\
\boldsymbol{J} & =\sigma \boldsymbol{E} & \begin{array}{c}
\text { Continuity } \\
\text { equation }
\end{array} \\
\frac{1}{\sigma} \nabla . \boldsymbol{J} & =\frac{\rho}{\epsilon_{0}} &
\end{array} \\
& -\frac{\partial \rho}{\partial t} \quad=\quad \frac{\sigma}{\epsilon_{0}} \rho \\
& \rho=\rho(0) e^{-\frac{\sigma_{0}}{\epsilon_{0}} t} \\
& \text { Div } \mathbf{J}=0 \left\lvert\, \begin{array}{c}
\text { steady current flow (as in a wire) } \\
\text { OR } \\
\mathbf{J}=0, \mathbf{E}=0 \text { (pure electrostatics) }
\end{array}\right. \\
& \text { In either case the excess charge is NOT in bulk } \\
& \text { But charge is conserved. } \\
& \text { Where does the excess charge go? } \\
& \text { From "bulk" to surface. } \\
& \text { How fast? } \\
& \text { For good metals like } \quad \mathrm{Cu}, \mathrm{Ag}, \mathrm{Au} \text {. } \\
& \sigma \sim 10^{7}-10^{8} \mathrm{~S} / \mathrm{m} \\
& \frac{\sigma}{\epsilon_{0}} \sim 10^{20} \\
& \text { time constant } \sim 10^{-20} \text { sec }
\end{aligned}
\]

You should now be able to justify the following :
- If no current is flowing in the conductor then \(\mathrm{E}=0\) inside
- All the excess charge resides on the surface of the conductor, even if there is a current flow.
- The conductor is an equipotential if \(\mathrm{J}=0\) (pure electrostatics)
- The electric field is normal to the surface (gradient is perpendicular to an equipotential)
- The surface charge density on a metal depends on the local radius of curvature.
- The electric field is strongest just outside sharp pointy edges.

Two spheres of radii \(R, r(R \gg r)\) are in contact so at the same potential. Which one has a larger E just outside? Why?
Think of a cone as made of a series of successively smaller spheres...where is the electric field strongest?

Generic Electrostatic Boundary conditions (normal and tangential components)

\[
\begin{aligned}
\oiint \boldsymbol{E} \cdot \boldsymbol{d} \boldsymbol{S} & =\frac{Q_{e n c}}{\epsilon_{0}} \\
E_{\perp \text { above }}-E_{\perp \text { below }} & =\frac{\sigma}{\epsilon_{0}}
\end{aligned}
\]

Normal component may be discontinuous
Surfaces may be finite.
The charge density may vary from place to place.
Surface is not necessarily equipotential

\[
\begin{array}{ccc}
\oint \boldsymbol{E} \cdot \boldsymbol{d} \boldsymbol{l} & =0 \\
E_{\| \text {above }}-E_{\| \text {below }} & =0
\end{array}
\]

But electrostatic potential V is always continuous at a surface. Reason: The discontinuity in \(E\) is finite. So \(V_{2}-V_{1}=-E . d l\) will go to zero as dl goes to zero

A scalar function \(V(\boldsymbol{r})\) satisfies \(\nabla^{2} V=0\)
Consider a sphere of radius \(R\) :integrate \(\nabla^{2} V\) over the volume
\[
\int_{v o l} \nabla \cdot(\nabla V) d \tau=\int_{\text {surface }} \nabla V \cdot d \boldsymbol{S} \quad \text { Gradient in spherical polar }
\]
\[
\begin{aligned}
& =\int\left[\boldsymbol{\epsilon}_{r} \frac{\partial V}{\partial r}+\boldsymbol{\epsilon}_{\theta} \frac{1}{r} \frac{\partial V}{\partial \theta}+\boldsymbol{\epsilon}_{\phi} \frac{1}{r \sin \theta} \frac{\partial V}{\partial \phi}\right] \cdot d \boldsymbol{S} \\
& =\int \frac{\partial V}{\partial r} R^{2} \sin \theta d \theta d \phi \quad \begin{array}{l}
\text { Only the radial component survives } \\
\text { because dS points radially outwards }
\end{array} \\
0 & =R^{2} \frac{\partial}{\partial r} \int_{\text {surface }} V(r, \theta, \phi) \sin \theta d \theta d \phi
\end{aligned}
\]

The average value \(\langle V(\theta, \phi)\rangle_{r}\) over a sphere is independent of \(r\). In the limit \(r \rightarrow 0\), we must have \(\langle V\rangle=V(0)\) So average value over a spherical surface \(=\) value at the center

Prove the 2D (using plane polar) and 1d cases as exercise.

There are no maxima or minima of \(V\) in a region where \(\nabla^{2} V=0\) But there can be saddle points


No stable equilibrium possible in purely electrostatic field (Earnshaw) All extremal values must occur at the boundary
\(V=\) const on ALL points on ALL boundaries \(\Rightarrow V\) is constant everywhere UNIQUENESS: There is only one possible solution of \(\nabla^{2} V=-\frac{\rho}{\epsilon_{0}}\) consistent with a given boundary condition

Two functions \(V_{1}\) and \(V_{2}\) satisfy \(\nabla^{2} V=-\frac{\rho}{\epsilon_{0}}\) With the same boundary conditions

Let \(\psi=V_{1}-V_{2} \quad\) Then \(\nabla^{2} \psi=0\) and \(\psi=0 \quad\) on ALL boundaries Implies \(\psi=0\) everywhere

Another way to prove this: consider the vector function: \(\psi \nabla \psi\)


If a "guess" satisfies the boundary condition then that MUST be the solution
\(\int_{\text {vol }}\left[\psi \nabla^{2} \psi+|\nabla \psi|^{2}\right] d \tau=0\)
\[
\int|\nabla \psi|^{2} d \tau \quad=\quad 0 \text { Possible only if } \psi=\text { constant=0 everywhere }
\]

\section*{Why is a metal cavity a "shield"?}

Arbitrary charges are outside the cavity (Q1...Qn)
Charges will be induced in the wall of the cavity.
But the wall remains an equipotential.
Inside the cavity V=0 is one possible solution satisfying the boundary conditions.

THAT IS THE UNIQUE SOLUTION.
What if the wall is not fixed at \(\mathrm{V}=0\) (i.e. floating)?
\(\mathrm{V}=\) constant is still correct, but the constant will depend on the charge distribution outside.

If charges are placed inside?
\(\nabla^{2} V=-\frac{\rho_{\text {in }}}{\epsilon_{0}}\)
\(V=0\) (boundary condition) irrespective of \(\rho_{\text {out }}\)


Floating conductor
Equal amounts +Q and -Q on inner and outer surfaces.

\section*{Why is a metal cavity a "shield" ?}


Floating conductor
Equal amounts \(+Q\) and \(-Q\) on inner and outer surfaces.

Special case: irregular cavity inside a sphere.

The charge density \(\sigma(\theta, \phi)\) is uniform
Surface is equipotential and \(\mathrm{E}=0\) inside. Use the boundary condition on Normal component of E and local charge density to prove the result.
\[
E(R)=\frac{Q}{4 \pi \epsilon_{0} R^{2}}
\]

Irrespective of the location of \(Q\) inside the cavity.

\section*{Electrostatic pressure on a conducting shell/ charged bubble}


Force on the element idS is due to E field created by all the other charges.

If \&s was treated in isolation


Outward Pressure=Force/area
The conducting surface simplifies the calculation. For an arbitrary surface

The difference must be due to the field created by all the other charges.

The force on IS is Charge in idS x Field due to all charges NOT in idS
it is more complex.....
\(\delta F=(\sigma \delta S) \frac{\sigma}{2 \epsilon_{0}}=\left(\frac{\sigma^{2}}{2 \epsilon_{0}}\right) \delta S\)

Electrostatics of conductors : uniqueness theorem 2 \& capacitance

If the charge on ALL the conductors is specified then the potential \(\mathrm{V}(\mathrm{x}, \mathrm{y}, \mathrm{z})\) is uniquely determined.

Notice that we are not specifying the charge distribution, only the total charge. That's the nontrivial content.

Suppose two distinct solutions exist \(U(x, y, z), V(x, y, z):\) Both must satisfy
\[
\int_{\text {surf } i}(-\nabla U) d \tau=\int_{\text {surf } i}(-\nabla V) d \tau=Q_{i}
\]
define \(\psi=U-V, \quad \&\) integrate \(\psi \nabla \psi\) over all surface
\(\sum_{i} \int_{\text {surf } i}(-\nabla \psi) \cdot d \boldsymbol{S}=\int_{\text {all vol }}\left[\nabla^{2} \psi+|\nabla \psi|^{2}\right] d \tau\) \(L H S=0: w h y ?\)
\[
\text { So } \int_{\text {all vol }}|\nabla \psi|^{2} d \tau=0 \quad=0
\]

Hence \(\quad U-V=0\)

U and V must give equipotentials On each conducting surface, but we do not claim that they are the same constant to start with.

Electrostatics of conductors : uniqueness theorem 2 \& capacitance

You are given the charge Q1, Q2....Qn On each conductor.

You are not told what \(\mathrm{V} 1, \mathrm{~V} 2 \ldots . . \mathrm{Vn}\) are.
What is the most generic statement you can make?
\[
\begin{aligned}
& Q_{i}=\sum_{j} C_{i j} V_{j} \\
& C_{i j}=C_{j i}
\end{aligned}
\]


The coefficients in this LINEAR relation are the formal definition of CAPACITANCE.

For a single object it reduces to the familiar relation : \(\mathbf{Q}=\mathbf{C V}\)
For an N -condcutor system the matrix is symmetric and can be inverted.
Try writing the energy of the system in matrix form as an exercise...

\section*{Where all do we come across Laplace's equation?}
1. Fluid flow : Incompressible, "inviscid", "irrotational"
flow of "DRY water", quite far from reality, still useful as a starting point
\[
\begin{aligned}
& (\rho=\text { const. } \quad \eta=0) \Rightarrow \nabla . v=0 \\
& \text { If } \nabla \times v=0 \text { then } v=\nabla \phi(\text { velocity potential }) \\
& \nabla^{2} \phi=0
\end{aligned}
\]
2. Heat conduction (Fourier), Diffusion equation (in steady state, time derivative \(=0\) )
\[
D \nabla^{2} \theta=\frac{\partial \theta}{\partial t}
\]
3. Electrostatic lensing :

Electron microscope, Ion trap, particle acceleration/beam steering, mass spectrometer Interesting differences from optical lensing:

Charged nature of particles,
Not possible to have focussing from all sides

Electrostatic 'quadrupole" lens : an example


\section*{Solving the Laplace equation : Image charge method}

Problem: A charge distribution and some boundary conditions are given. The usual boundary conditions are fixed potentials over some surfaces.
Solve for \(V(r)\) in a certain region.
A "trick" works for some (!! not all !!) problems.
STEP 1: put some point charges in the regions NOT part of the region where you need to solve for the potential.

STEP 2: Try to arrange these external charges, so that the external + given charges together produce the desired potential at the boundaries. Forget all else!

STEP 3: Calculate the total potential in the certain (given) region using all the charges in the problem + external charges.

STEP 4: The total field/potential produced by the ALL the charges is the solution to the problem. The extra charges are called Image charges.

Image charge method : point charge near a conducting grounded plane

\(z=0\)

\section*{PROBLEM:}

A point charge \(+Q\) is kept at \(\left(0,0, z_{0}\right)\)
\(z=0\) is a grounded conducting sheet.
What is \(V(x, y, z)\) for \(z>0\)
Subject to boundary conditions:
\(V=0\) at \(z=0\)
\(V \rightarrow 0\) as \(\mathrm{x}, \mathrm{y}, \mathrm{z} \rightarrow \infty\)

\section*{SOLUTION:}

Put an extra charge \(-Q\) at \(\left(0,0,-z_{0}\right)\)
The potential due to both is
\[
V(x, y, z)=\frac{Q}{4 \pi \epsilon_{0}}\left[\frac{1}{\sqrt{x^{2}+y^{2}+\left(z-z_{0}\right)^{2}}}-\frac{1}{\sqrt{x^{2}+y^{2}+\left(z+z_{0}\right)^{2}}}\right]
\]

For \(\mathrm{z}=0\), the two terms cancel giving \(\mathrm{V}=0\). This must be the solution (uniqueness). We can now calculate the force between the charge \(+Q\), induced surface charge at every point etc.

Image charge method : point charge near a conducting grounded plane


To solve this we need three image charges:
-Q at \((\mathrm{a},-\mathrm{b})\)
\(+Q\) at \((-a,-b)\)
\(-Q\) at \((-a, b)\)

This problem can be solved with a finite number of images if
\(\theta \times\) integer \(=\pi\)

There is no generic method! It is a combination of guess and some calculation.....

Image charge method : point charge near a conducting grounded sphere


Solution wanted outside the sphere only

The distance of an arbitrary point on the surface from the charges \(Q\) and \(Q^{\prime}\)
\[
\begin{aligned}
d^{\prime 2} & =r^{2}+a^{2}-2 a r \cos \theta \\
d^{2} & =r^{2}+b^{2}-2 b r \cos \theta
\end{aligned}
\]
can we adjust a and Q' such that
\[
\frac{Q}{d}-\frac{Q^{\prime}}{d^{\prime}}=0 \quad \text { for all } \theta ?
\]
\[
\frac{r^{2}+a^{2}-2 a r \cos \theta}{Q^{\prime 2}}=\frac{r^{2}+b^{2}-2 b r \cos \theta}{Q^{2}}
\]
\[
a=\frac{r^{2}}{b}
\]

Equate the terms independent of \(\cos \theta\) and coefficient of \(\cos \theta\) on both sides
\(\frac{Q^{\prime}}{Q}=\frac{r}{b}\)

Image charge method : off centred point charge inside grounded sphere


Hollow sphere of radius \(r\) is kept at \(V=0\) Inside the sphere there is a charge \(+Q\) placed at a distance a from the center.

What is the potential inside the sphere?
Notice the "conjugate" nature of this problem with the last one.

This is a characteristic of "image charge problems".

How would you adapt the image charge method for a case where the spherical surface is at a potential \(\vee \neq 0\) ?
\[
\begin{aligned}
& \text { 1D:(trivial!) } \\
& \frac{d^{2} V}{d x^{2}}=0 \Rightarrow V=A x+B
\end{aligned}
\]

Cannot be anything more complicated.
2D :(cartesian)

In simplest (very few!) cases
\[
\frac{\partial^{2} V}{\partial x^{2}}+\frac{\partial^{2} V}{\partial y^{2}}=0
\] separation of variable will work.
\(A, B, C, D\) are chosen
\[
\text { If } V(x, y)=X(x) Y(y) \text { then }
\] to match the given boundary conditions.
\[
\frac{1}{X} \frac{d^{2} X}{d x^{2}}+\frac{1}{Y} \frac{d^{2} Y}{d y^{2}}=0
\]

Role of \(x, y\) can be Interchanged.
\[
\frac{1}{X} \frac{d^{2} X}{d x^{2}}=-\frac{1}{Y} \frac{d^{2} Y}{d y^{2}}=k^{2}(\text { const })
\]

X may be exponential \&
Y may be sinusoidal.

Solution \(=\) sinusoidal type \(\times\) exponential decay or growth
\[
V(x, y)=\sum_{\text {allowed } k}\left(A_{k} \cos k x+B_{k} \sin k x\right)\left(C_{k} e^{k y}+D_{k} e^{-k y}\right)
\]

Basic idea: Take any analytic complex function (eg. \(z^{2}, \sin z, e^{-2}\) )
\[
F(z)=u(x, y)+i v(x, y)
\]

\section*{Both \(u(x, y)\) and \(v(x, y)\) satisfy 2D Laplace equation}

By intuition/guess/imagination make \(u(x, y)\) or \(v(x, y)\) satisfy the boundary conditions only. Uniqueness theorem guarantees that the guess is THE solution.

In reality, very few problems can be solved by separation of variables.
Quite a few can be done by the complex number method - but in 2D only.
Particularly useful for solutions in near corners, slits, edges, quadrupoles.
\[
\begin{aligned}
& F(z)=i \ln z^{2} \\
& u(x, 0)=0 \\
& u(0, y)=\pi
\end{aligned}
\]

Use this fact and scale the function as \(F(z)=-\left(V_{2}-V_{1}\right) i \ln z^{2}+V_{1}\)


Notice that separation of variable doesn't work here.
How will you modify the solution if the two plates are inclined at an angle \(\alpha\) ?


A semi-infinite plate occupies the region \(x>0\) in the \(x y\) plane.

The plate is kept at \(\mathrm{V}=\) const.
The problem shows how to handle the field near the edge of a flat thin plate.
\[
F(z)=z^{1 / 2}
\]
\[
=\rho^{1 / 2}\left(\cos \frac{\theta}{2}+i \sin \frac{\theta}{2}\right)
\]
\[
=\left[\frac{\left(x^{2}+y^{2}\right)^{1 / 2}+x}{2}\right]^{1 / 2}+i\left[\frac{\left(x^{2}+y^{2}\right)^{1 / 2}-x}{2}\right]^{1 / 2}
\]

Notice \(v(x, 0)=0\) if \(x>0\)
\(v(x, y)+V_{0}\) satisfies required boundary condition

Solving Laplace's equation (2D : Using complex numbers : edge)


From: Feynman Lectures vol. 2

The figure is drawn assuming the fixed potential is zero. It can of course be shifted by any amount

\section*{Plate}

Solving Laplace's equation (2D : Using complex numbers : slit)


Two semi-infinite platea occupies the region \([-a<x<a]\) in the \(x z\) plane.

The plates are kept at VL and VR
How to handle the field in a slit between equipotential plates?

The solution of this problem requires a transformation of the complex variables called "conformal transform". See: Pipes \& Harvill

You cannot superpose two plates at fixed potentials to get a slit. Why?
\(v(x, y)=V_{L}+\frac{V_{R}-V_{L}}{\pi}\left[\arcsin \frac{1}{2}\left(\sqrt{(x / a+1)^{2}+(y / a)^{2}}-\sqrt{(x / a-1)^{2}+(y / a)^{2}}\right)+\frac{\pi}{2}\right]\)
\[
\nabla^{2} V=\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial V}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} V}{\partial \theta^{2}}=0
\]

Try: \(V=R(r) e^{i m \theta}\)
Why not \(e^{m \theta}\) ?
This gives:
\(r^{2} \frac{d^{2} R}{d r^{2}}+r \frac{d R}{d r}-m^{2} R=0\)
trial solution \(R=A r^{n}\) gives : \(n= \pm m\), so
\(V(r, \theta)=\sum_{m}\left(A_{m} r^{m}+\frac{B_{m}}{r^{m}}\right) e^{i m \theta}\)

Why should m be an integer?

What type of problems can we solve with this form?

Values given on a circle.
Solution inside should not have \(1 / r\) type solution
Solution outside (till infinity) should not have \(r\) type solution. 1
Use Fourier analysis to find the coeffcients.

Solving Laplace's equation (3D : Spherical polar)
\(\nabla^{2} V=\frac{1}{r^{2}} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial V}{\partial r}\right)+\frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial V}{\partial \theta}\right)+\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2} V}{\partial \phi^{2}}\)
With no \(\phi\) dependence we try: \(V(r, \theta)=R(r) P(\theta)\)
\(\frac{1}{R} \frac{\partial}{\partial r}\left(r^{2} \frac{\partial R}{\partial r}\right)=-\frac{1}{P} \frac{1}{r^{2} \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial V}{\partial \theta}\right)=l(l+1)\)

The radial solution
\[
r^{2} \frac{d^{2} R}{d r^{2}}+2 \mathrm{r} \frac{d R}{d r}-l(l+1) R=0
\]
try \(R=A r^{n}\)
\(R=A r^{l}+\frac{B}{r^{l+1}}\)

Notice the utility of wiritng the seperation constant in the \(1(1+1)\) form

\section*{Solving Laplace's equation (3D : Spherical polar)}

The angular part:
\(\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta \frac{\partial P}{\partial \theta}\right)+l(l+1) P=0\)

Polynomial solutions worked in the examples before this, but would not work in this case. Why?
substitute \(x=\cos \theta\)
\(\left(1-x^{2}\right) \frac{d^{2} P}{d x^{2}}-2 x \frac{d P}{d x}+l(l+1) P=0\)
try the series: \(P=\sum_{0}^{\infty} a_{n} x^{n}\) : this gives
\(\left(1-x^{2}\right) \sum n(n-1) a_{n} x^{n-2}-2 x \sum n a_{n} x^{n-1}+l(l+1) \sum a_{n} x^{n}=0\)
\[
2 \mathrm{a}_{2}+l(l+1) a_{0}=0
\]
3.2. \(a_{3}-2 \mathrm{a}_{1}+l(l+1) a_{1}=0\)
\(a_{n+2}=-\frac{(l-n)(l+n+1)}{(n+2)(n+1)} a_{n}\)
a 0 and a1 can be arbitrarily chosen
If \(I\) is an integer, then the series will terminate at \(\mathrm{n}=1\)

Odd and even powers do not mix in this recurrence relation

\section*{Solving Laplace's equation (3D : Spherical polar)}
\[
P(x)=a_{0} \sum(\text { even powers of } x)+a_{1} \sum(\text { odd powers of } x)
\]

So construct each polynomial using the recurrence relation
\[
\begin{aligned}
& P_{0}(x)=1 \\
& P_{1}(x)=x \\
& P_{2}(x)=\frac{1}{2}\left(3 x^{2}-1\right) \\
& P_{3}(x)=\frac{1}{2}\left(5 x^{3}-3 x\right)
\end{aligned}
\]

\section*{Legendre Polynomials:}
\[
\int_{-1}^{1} P_{m}(x) P_{n}(x) d x=\frac{2}{2 \mathrm{n}+1} \delta_{n m}
\]

Use orthogonality to find expansion co-effs ...

Solutions of a general class of diffn equations have this orthogonality property - called
"Sturm-Liuville" diffn eqn

Values given on a sphere.
Solution inside should not have \(1 / r\) type solution
Solution outside (till infinity) should not have \(r\) type solution.
Use expansion in Legendre Polynomials to find the coeffcients.
See the worked examples of Griffiths...section 3.3

Solving Laplace's equation (3D : Spherical polar)

\section*{Legendre Polynomials}


To generate the successive \(P_{l}(x)\) use the Rodrigue's formula:
\(P_{l}(x)=\frac{1}{2^{l} l!} \frac{d^{l}}{d x^{l}}\left(x^{2}-1\right)^{l}\)

\section*{Work and Energy in electrostatics}

Conservative field : Total energy of a particle is conserved.
KE+PE is conserved. Or equivalently

Work done in moving a particle very slowly from one point to another is path independent.

A potential energy function exists
Gravitational potential:
Apple falling from a tree Earth going round the sun... Trajectory of a particle...
The force is derivable from a scalar potential

Curl of the Force field is zero.

Why do we need to say more?
The answer to this is not within "electrostatics"....the need really comes When we deal with \(E, B\) and moving charges.

Work done on the charge
\[
\begin{aligned}
W & =-q \int_{a}^{b} \boldsymbol{E} . \boldsymbol{d r} \quad[\text { always true }] \\
& =q\left(V_{b}-V_{a}\right) \quad[\text { only if } \nabla \times \boldsymbol{E}=0]
\end{aligned}
\]

SI unit is Joule.
Useful unit is electron-Volt
Work needed to move one electron through 1 volt \(\mathrm{q}=-1.6 \times 10^{-19} \mathrm{C}\)

How do we build up a configuration of charges?
Bring the first charge : No work done
Bring the second charge from infinity to desried position : calculate work done
Bring next one. Calculate the work done due to the presence of the previous TWO

\section*{Work and Energy in electrostatics}
\[
W=\frac{1}{4 \pi \epsilon_{0}} \sum_{j<i} \frac{q_{i} q_{j}}{\left|\boldsymbol{r}_{i}-\boldsymbol{r}_{j}\right|}=\frac{1}{2} \sum_{i} q_{i} \frac{1}{4 \pi \epsilon_{0}} \sum_{j} \frac{q_{j}}{\left|\boldsymbol{r}_{i}-\boldsymbol{r}_{j}\right|}
\]

Potential at location of charge i

Now suppose this is a continous distribution : it means we are saying the following
\[
\begin{aligned}
\sum_{i} q_{i}(\ldots . .) & \int_{\text {all space }} \rho(\boldsymbol{r}) d \tau(\ldots .) \quad \text { When can this breakdown? } \\
\frac{1}{2} \int \rho V d \tau & =\frac{\epsilon_{0}}{2} \int(\nabla . \boldsymbol{E}) V d \tau \\
& =\frac{\epsilon_{0}}{2} \int\left[\nabla \cdot \left(\widehat{\boldsymbol{E} V)-\boldsymbol{E}(-\nabla V)] d \tau} \begin{array}{c}
\text { Convert to surface integral } \\
\text { Take surface at infinity } \\
\text { Should go to zero }
\end{array}\right.\right. \\
& =\frac{\epsilon_{0}}{2} \int[\boldsymbol{E} . \boldsymbol{E}] d \tau
\end{aligned}
\]
\[
\frac{\epsilon_{0}}{2} \int_{0}^{\infty}\left(\frac{q}{4 \pi \epsilon_{0} r^{2}}\right) d \tau
\]

The closest we can try :
Assume that a point charge is a uniform sphere of some radius R .

The integral for field energy will then converge.

\[
\begin{aligned}
& \rho=\frac{Q}{\frac{4}{3} \pi r^{3}} \\
& E_{\text {in }}=\frac{Q}{4 \pi \epsilon_{0} R^{2}} \frac{r}{R} \quad(r<R) \\
& E_{\text {out }}=\frac{Q}{4 \pi \epsilon_{0} r^{2}} \quad(r>R) \\
& \left(\frac{\epsilon_{0}}{2}\right)\left[\int_{0}^{R} E_{\text {in }}^{2} d \tau+\int_{R}^{\infty} E_{\text {out }}^{2} d \tau\right] \quad \text { will converge }
\end{aligned}
\]

Within classical electromagnetism it is not possible to resolve this problem.

We have to accept that the concept of a point charge has some limitations

Dielectric materials:
Field of a polarised object at a large distance
Multipole expansion of scalar potential
Polar and cartesian expressions for dipole, quadrupole etc Atomic and molecular origin of the dipole moment Equivalent charge distribution
Force and torque on a dipole
Definition of the E D P vectors and boundary conditions Interface of two dielectrics, sphere in an uniform field
Energy contained in Electric fields with dielectrics present

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Quantitatively this means : With what power law does it fall off ....inverse square, cube, fourth ?
Often the charge is limited to a small area.
Answer:
\[
V(\vec{r})=\frac{1}{4 \pi \epsilon_{0}} \int d^{3} \overrightarrow{r^{\prime}} \frac{\rho\left(\overrightarrow{r^{\prime}}\right)}{\left|\vec{r}-\overrightarrow{r^{\prime}}\right|}
\]

In many cases \(r \gg r^{\prime}\)
So expand in a power series in \(\frac{r^{\prime}}{r}\)
From the figure :
\[
\begin{aligned}
\frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|} & =\left[r^{2}+r^{\prime 2}-2 r r^{\prime} \cos \theta\right]^{-\frac{1}{2}} \\
& =\frac{1}{r}\left[1-\left\{2 \frac{r^{\prime}}{r} \cos \theta-\left(\frac{r^{\prime}}{r}\right)^{2}\right\}\right]^{-\frac{1}{2}} \\
& =\frac{1}{r} \sum_{l=0}^{\infty}\left(\frac{r^{\prime}}{r}\right)^{l} P_{l}(\cos \theta)
\end{aligned}
\]

Binomial expansion .....
\[
\begin{aligned}
& (1-x)^{-\frac{1}{2}} \\
= & 1+\frac{1}{2} x+\frac{3}{8} x^{2}+\frac{5}{16} x^{3}+\frac{35}{128} x^{4} \ldots
\end{aligned}
\]
\[
\begin{aligned}
V(P)= & \frac{1}{4 \pi \epsilon_{0}} \sum_{l=0}^{\infty} \frac{1}{r^{l+1}} \int d^{3} \vec{r}^{\prime}\left[\rho\left(\overrightarrow{r^{\prime}}\right) r^{\prime l} P_{l}(\cos \theta)\right] \\
= & \frac{1}{4 \pi \epsilon_{0}}\left[\frac{1}{r} \int d^{3} \vec{r}^{\prime} \rho\left(\vec{r}^{\prime}\right)+\right. \\
& \frac{1}{r^{2}} \int d^{3} \vec{r}^{\prime} r^{\prime} \cos \theta \rho\left(\vec{r}^{\prime}\right)+ \\
& \left.\frac{1}{r^{3}} \int d^{3} \vec{r}^{\prime}\left(r^{\prime}\right)^{2} \frac{1}{2}\left(3 \cos ^{2} \theta-1\right) \rho\left(\overrightarrow{r^{\prime}}\right)+\ldots\right]
\end{aligned}
\]

If the total charge is zero:
Dipole term dominates.
If that is also zero
Quadrupole dominates
suppose
\[
\rho\left(\vec{r}^{\prime}\right)=q \delta\left(\overrightarrow{r^{\prime}}-\vec{a}\right)-q \delta\left(\overrightarrow{r^{\prime}}+\vec{a}\right)
\]
how will the dipole integral look?
can write this as
\[
\begin{aligned}
V_{\text {dipole }} & =\frac{1}{4 \pi \epsilon_{0}} \frac{\vec{p} \cdot \overrightarrow{\hat{r}}}{r^{2}} \\
\text { with } \vec{p} & =\sum q_{i} \vec{r}_{i}^{\prime}
\end{aligned}
\]

and some other equivalent forms...

Choice of the co-ordinate system and origin in multipole expansion
We could have done the expansion in a more cartesian way...
\(\frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|}=\left[r^{2}+r^{\prime 2}-2 \vec{r} \cdot \vec{r}^{\prime}\right]^{-\frac{1}{2}}\)
This would have given succeessive terms like....
\[
\begin{aligned}
& V_{\text {mono }}=\frac{1}{4 \pi \epsilon_{0}} \frac{Q_{\text {total }}}{r} \\
& V_{\text {dip }}=\frac{1}{4 \pi \epsilon_{0}} \frac{\sum \hat{r}_{i} p_{i}}{r^{2}} \\
& V_{\text {quad }}=\frac{1}{4 \pi \epsilon_{0}} \frac{1}{2} \frac{\sum_{i j} \hat{r}_{i} \hat{r}_{j} Q_{i j}}{r^{3}}
\end{aligned}
\]
\[
\begin{aligned}
p_{i} & =\int d^{3} \vec{r}^{\prime} r_{i}^{\prime} \rho\left(r^{\prime}\right) \\
Q_{i j} & =\int d^{3} \vec{r}^{\prime}\left(r_{i}^{\prime} r_{j}^{\prime}-r^{\prime 2} \delta_{i j}\right) \rho\left(r^{\prime}\right)
\end{aligned}
\]

Dipole moment is a vector Quadrupole moment is a tensor

The lowest non-vanishing moment is independent of the choice of the origin.
The higher moments are NOT necessarily so.
So if the total charge (monopole) is zero then dipole term is origin-independent. If the dipole also vanishes then quadrupole is origin independent. (Prove it!)

Dipole is more common in electronic charge distributions.
Nucleii often have quadrupole moments.
Earth's gravitational potential has a significant quadrupole component.

\section*{Response of atoms and molecules to an electric field}


Fixed


\(Q\) is the total charge
\(A\) is the atomic radius
\(p=Q d\) (induced moment)

Electron cloud is an uniformly charged sphere..(assume)
Force on the nucleus due to displaced electron cloud = External force on nucleus
\[
E=\frac{1}{4 \pi \epsilon_{0}} \frac{Q d}{a^{3}} \quad \text { hence } \quad \vec{p}=4 \pi \epsilon_{0} a^{3} \vec{E}
\]
\[
\Delta \sim 4 \pi \epsilon_{0} \times 10^{-30} \quad \text { in SI }
\]

Atomic polarizability Small for inert gases
Large for atoms with partially filled outer shell
Estimated values and observed value agree (order of magnitude)

\(6.2 \times 10^{-30}\) Coul.mt \((1.85 \mathrm{D})\)


No dipole moment

SI unit \(=\) Coul-mt.
1 Debye unit (historical but useful) Dipole moment of \(10^{-10}\) esu of charge separated by 1 angstrom Useful for molecular scale since Electron charge is \(4.8 \times 10^{-10} \mathrm{esu}\)

Electron distribution in the bonds can give rise to built in dipole moment


Induced dipole moment and electric field are not necessarily in the same direction for a molecule. Since the bonds do not shift uniformly in all directions...."easy" and "hard" directions... \(P\) and \(E\) are related by a matrix/tensor

\section*{Force and torque on a dipole}

Potential Energy and force

\[
\begin{aligned}
& \vec{p}=q \delta \vec{r} \\
& U_{d i p}=-\vec{p} \cdot \vec{E} \\
& \vec{F}_{\text {dip }}=(\vec{p} \cdot \vec{\nabla}) \vec{E}
\end{aligned}
\]

Torque


Although the E field is different at two sites, the difference in the final expression would be second order.....

Now we can calculate the interaction force between two dipoles....easily!
If we have two dipoles...the \(E\) field of the first will act on the second and vice versa,

\section*{Potential of an extended distribution of dipoles}
\[
\begin{aligned}
& V(X)=\frac{1}{4 \pi \epsilon_{0}} \int d^{3} \vec{r}^{\prime} \vec{P} \cdot \frac{\vec{r}-\vec{r}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}} \\
& \nabla_{r^{\prime}} \frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|}=\frac{\vec{r}-\vec{r}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}} \quad \begin{array}{l}
\text { Prove this by writing out } \\
\text { in (x-x')..... }
\end{array}
\end{aligned}
\]

\section*{Hence}
\[
\begin{aligned}
V(X) & =\frac{1}{4 \pi \epsilon_{0}} \int d^{3} \vec{r}^{\prime}\left[\nabla \frac{\vec{P}}{\left|\vec{r}-\vec{r}^{\prime}\right|}-\frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|} \nabla \cdot \vec{P}\right] \\
& =\frac{1}{4 \pi \epsilon_{0}}\left[\int \frac{d \vec{S}^{\prime} \cdot \vec{P}}{\left|\vec{r}-\vec{r}^{\prime}\right|}-\int d^{3} \vec{r}^{\prime} \frac{\nabla \cdot \vec{P}}{\left|\vec{r}-\vec{r}^{\prime}\right|}\right]
\end{aligned}
\]

Surface charge Volume charge
\(\sigma=\vec{P} . \hat{n}\)
\(\rho=-\nabla \cdot \vec{P}\)
dipole distribution to be integrated over

Here integration and differentiation are w.r.t. primed co-ordinates

\section*{Linear Dielectrics: E P D vectors}

Linear dielectric means : Induced dipole moment ( \(\mathbf{P}\) ) is proportional to the electric field. Hence:
\[
\begin{aligned}
& \nabla \cdot \vec{E}=\frac{\rho_{\text {TOTAL }}}{\epsilon_{0}} \\
& \nabla . \epsilon_{0} \vec{E}=\rho_{\text {free }}+\rho_{\text {pol }} \quad\left(\text { since } \quad \rho_{\text {pol }}=-\nabla . \vec{P}\right) \\
& \nabla \cdot\left[\epsilon_{0} \vec{E}+\vec{P}\right]=\rho_{\text {free }} \\
& \text { Use the proprotionality of } \vec{P} \text { with } \vec{E}: \quad \vec{P}=\epsilon_{0} \chi \vec{E} \\
& \epsilon_{0}(1+\chi) \vec{E}=\epsilon \vec{E}=\vec{D} \quad \text { Historically called electric displacement vector: } \\
& \text { - Microscopic mechanism was not known then. } \\
& \nabla \cdot \vec{D}=\rho_{\text {free }} \\
& \text { Quantities like } \mathrm{D}, \varepsilon \text { can only be defined in an average sense. } \\
& \nabla \cdot \vec{E}=\frac{\rho_{\text {free }}}{\epsilon} \\
& \text { !! One cannot talk about } \mathrm{D} \text { or } \varepsilon \text { inside an atom!! } \\
& \nabla \times \vec{E}=0 \\
& \text { Since curl } \mathbf{E}=0 \text {, a scalar potential is still possible. } \\
& \text { But the "source" of this potential is reduced by a factor. } \\
& \text { Hence the scalar potential } \mathrm{V} \text { is also reduced by that factor. }
\end{aligned}
\]

Linear Dielectrics: EP D vectors: Boundary conditions and related problems
\(\begin{array}{cc}E_{1} \\ D_{1}= \\ \varepsilon_{1} E_{1} \\ & E_{2} \\ & D_{\perp}=\varepsilon_{2} E_{2} \\ & E_{11} \text { is continuous }\end{array}\)
No FREE CHARGE

Point charge \(q\) is placed at \((0,0, d)\) as shown near an interface of two dielectrics.
Q: What is the potential everywhere?
For \(z>0\) : (region 2)
Image charge q' at ( \(0,0,-\mathrm{d}\) )
For \(z<0\) : (region 1)
Charge q" at ( \(0,0, \mathrm{~d}\) )
Write the potential, then the electric field. Two independent equations by matching the normal and tangential components at the boundary.


\section*{Linear Dielectrics : A uniformly polarised sphere}

Uniformly polarised sphere : (no external field)
Note the lines of force (Electric field):
Points in the opposite direction inside the sphere.
\[
V(r, \theta)= \begin{cases}\sum_{l=0}^{\infty} A_{l} r^{l} P_{l}(\cos \theta) & (0<r \leqslant R) \\ \sum_{l=0}^{\infty} \frac{B_{l}}{r^{l+1}} P_{l}(\cos \theta) & (r \geqslant R)\end{cases}
\]

Use boundary conditions at \(r=R\) and
\(V\) should be well behaved at small and large \(r\)...
\(V(r=R)\) should match
E should have a discontinuity due to surface charge Equate the coefficient of each Legendre polynomial
\[
V(r, \theta)=\left\{\begin{array}{lc}
\frac{P}{3 \epsilon_{0}} r \cos \theta & (0<r \leqslant R) \\
\frac{P}{3 \epsilon_{0}} \frac{R^{3}}{r^{2}} \cos \theta & (r \geqslant R)
\end{array}\right.
\]


Surface charge : \(\sigma_{b}=\vec{P} \cdot \hat{n}=P \cos \theta\)
Volume charge : \(\rho_{b}=-\nabla . \vec{P}=0\)

Looks like the field of a single (pure) dipole at \(\mathrm{r}=0\)

The field is CONSTANT inside

Linear Dielectrics : A dielectric sphere in an uniform field
Uniform field means far from the sphere \(E=E_{0}\) set externally
\[
V(r, \theta)= \begin{cases}\sum_{l=0}^{\infty} A_{l} r^{l} P_{l}(\cos \theta) & (0<r \leqslant R) \\ \sum_{l=0}^{\infty}\left[B_{l} r^{l}+\frac{C_{l}}{r^{l+1}}\right] P_{l}(\cos \theta) & (r \geqslant R)\end{cases}
\]
\[
-\left.\epsilon \frac{\partial V_{\mathrm{in}}}{\partial r}\right|_{r=R}=-\left.\epsilon_{0} \frac{\partial V_{\text {out }}}{\partial r}\right|_{r=R} \quad \begin{aligned}
& \text { Normal component of } \mathrm{D} \text { is } \\
& \text { continuous }
\end{aligned}
\]
\[
-\frac{1}{R} \frac{\partial V_{\mathrm{in}}}{\partial \theta}=-\frac{1}{R} \frac{\partial V_{\mathrm{out}}}{\partial \theta}
\]
Tangential component of E is continuous
\[
V_{\mathrm{in}}=\left(\frac{3}{2+\epsilon / \epsilon_{0}}\right) E_{0} r \cos \theta
\]
\[
V_{\text {out }}=-E_{0} r \cos \theta+\left(\frac{\epsilon-\epsilon_{0}}{\epsilon+2 \epsilon_{0}}\right) E_{0} \frac{R^{3}}{r^{2}} \cos \theta
\]
What would \(\epsilon \rightarrow \infty\) physically mean?
If a spherical cavity is dug out in a large slab?

Notice how the orthogonality of Legendre polynomials is crucial to solving these problems

Linear Dielectrics : A capacitor with a dielectric slab

\(-\sigma_{f}\)
\[
\frac{V}{d}=E=\frac{\sigma_{f}}{\epsilon_{0}}
\]

For the same charge on the metal plate, the voltage developed is now smaller.
\(\mathrm{C}=\mathrm{Q} / \mathrm{V}\) thus increases by a factor of \(1+\chi\) The energy stored in the field also increases by the same factor if the same voltage and hence the same \(\mathrm{E}=\mathrm{V} / \mathrm{d}\) is established in the capacitor.

\[
\begin{array}{rlll}
\frac{V}{d} & = & E & =\frac{\sigma_{f}-\sigma_{b}}{\epsilon_{0}} \\
\sigma_{b} & = & \vec{P} \cdot \hat{n} & =\epsilon_{0} \chi E \\
& \text { so } \quad E & =\frac{\sigma_{b}}{\epsilon_{0} \chi} \\
\sigma_{f}-\sigma_{b} & =\frac{\sigma_{f}}{1+\chi}
\end{array}
\]
\[
\frac{V}{d}=\quad E=\frac{\sigma_{f}}{\epsilon_{0}(1+\chi)}
\]

Magnetostatics

\section*{Magnetostatics: Field due to steady currents}

Magnetic fields are created by:
1. Charges in motion
2. Intrinsic spin (dipole) moments of some elementary particles.

The field created by a single moving charge is not very easy to write down! Historically steady currents were understood first.

Maxwell's equation tell us the field created by currents (moving charges).
The response of a particle to the fields ( \(E\) and \(B\) ) is an independent input.
\[
\begin{aligned}
& \qquad \vec{F}=q(\vec{E}+\vec{v} \times \vec{B}) \\
& \nabla \vec{B}=0 \\
& \nabla \times \vec{B}=\mu_{0} \vec{J}+\epsilon_{0} \mu_{0} \frac{\partial \vec{E}}{\partial t} \\
& \begin{array}{l}
\text { Lorentz Force Law is } \\
\text { Not derivable } \\
\text { from Maxwell's equation }
\end{array} \\
& \begin{array}{l}
\text { Historically these equations } \\
\text { were written later. }
\end{array} \\
& \begin{array}{l}
\text { Zero divergence means : } \\
\text { No analogue of charge as in } \\
\text { electrostatics. }
\end{array} \\
& \mu_{0}=4 \pi \times 10^{-7} \text { Henry. }^{-1} \\
& \text { specifies a vector field fully if } \\
& \text { suitable boundary conditions } \\
& \text { are also given. }
\end{aligned}
\]


Magnets deflect a "compass" and so does a nearby current carrying wire. So a current carrying wire must be creating a magnetic field. (Oersted, Ampere)

\[
\overrightarrow{d B}=\frac{\mu_{0}}{4 \pi} \frac{I \vec{d} l \times \vec{r}}{r^{3}}
\]

Then integrate over the wire to find the full field.

Common geometries..... Straight lines, coils etc.

Field from a wire segment, loops, coils, toriods etc

\[
\begin{aligned}
& R=r \cos \theta \\
& l=R \tan \theta \\
& \delta l=R \sec ^{2} \theta \delta \theta
\end{aligned}
\]
\[
\begin{aligned}
\delta B & =\frac{\mu_{0}}{4 \pi} I \frac{\overrightarrow{\delta l} \times \vec{r}}{r^{3}} \\
& =\frac{\mu_{0} I}{4 \pi} \cdot \frac{R \sec ^{2} \theta \delta \theta \cdot \sin (\pi / 2-\theta) \cdot R \sec \theta}{R^{3} \sec ^{3} \theta} \\
B & =\frac{\mu_{0} I}{4 \pi} \cdot \int_{\theta}^{\theta} \cos \theta d \theta=\frac{\mu_{0} I}{4 \pi}\left(\sin \theta_{2}-\sin \theta_{1}\right)
\end{aligned}
\]

Infinite wire \(\left\{\begin{array}{l}\theta_{1}=-\pi / 2 \\ \theta_{2}=\pi / 2\end{array}\right.\).
\(B=\frac{\mu_{0} I}{2 \pi R} \cdot\) (Can use symmetry arguments)

Field from a wire segment, loops, coils, toriods etc

\[
\begin{aligned}
& r^{2}=R^{2}+z^{2} \\
& \sin \varphi=R / r \\
& \delta l=R \delta \theta
\end{aligned}
\]
\[
\begin{aligned}
\delta B & =\frac{\mu_{0}}{4 \pi} \cdot I(R \delta \theta) \cdot \frac{1}{R^{2}+z^{2}} \cdot \frac{R}{r} \\
B & =\frac{\mu_{0} I}{2} \cdot \frac{R^{2}}{\left(R^{2}+z^{2}\right)^{3 / 2}}
\end{aligned}
\]

Integrate over \(z\) for a finite coil.


Using the integral form in symmetrical cases: long wire, coil, full toroid
\[
\oint \vec{B} \cdot \overrightarrow{d l}=\mu_{0} I_{e n c} \quad \text { same as } \quad \nabla \times \vec{B}=\mu_{0} \vec{J}
\]

Infinitely long wire

\[
\begin{array}{ll}
B_{Q} 2 \pi r=\mu_{0} I & B_{r}=0 \quad(0, B=0) \\
B_{\phi}=\frac{\mu_{0} I}{2 \pi r} & B_{2}=0
\end{array}
\]

Infinitely long coil

\[
\begin{aligned}
& B L=\mu_{0} N I \\
& B=\mu_{0} \frac{N}{L} I
\end{aligned}
\]

Using the integral form in symmetrical cases: long wire, coil, full toroid


The magnetic vector potential : the formal solution
\[
\begin{aligned}
\nabla \times \vec{B} & =\mu_{0} \vec{J} \\
\nabla \times(\nabla \times \vec{A}) & =\mu_{0} \vec{J} \\
\nabla(\nabla \cdot \vec{A})-\nabla^{2} \vec{A} & =\mu_{0} \vec{J}
\end{aligned}
\]

The choice \(\nabla \cdot \vec{A}=0\) is called a gauge choice

Since \(\nabla \cdot \vec{B}=0\) we can write \(\vec{B}=\nabla \times \vec{A}\)
Does it make thing simpler?

There can be other possible choices.
For each gauge \(A\) and \(V\) will be different But they will give the same \(\mathbf{E}\) and \(\mathbf{B}\).
\[
\nabla^{2} \vec{A}=-\mu_{0} \vec{J}
\]
is like three Poisson's equation \(\left(\nabla^{2} V=-\frac{\rho}{\epsilon_{0}}\right)\) put together
\(\operatorname{Can} \vec{A}\) the vector potential.
\(\vec{A}\) has wo simple interpretation like potential energy

The magnetic vector potential : the formal solution
\(A(\vec{r})=\frac{\mu_{0}}{4 \pi} \int \frac{\vec{J}\left(\vec{r}^{\prime}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|} d^{3} \vec{r}^{\prime}\)
\(\vec{B}=\nabla \times \overrightarrow{(\vec{r})}=\nabla \times \frac{\mu_{0}}{4 \pi} \int \frac{\vec{J}\left(\vec{r}^{\prime}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|} d^{3} \vec{r}^{\prime}\)

The curl is to be taken w.r.t. \(\vec{r}\) The integration is w.r.t. \(\vec{r}^{\prime}\)
what is \(\nabla \times \frac{\vec{J}\left(\vec{r}^{\prime}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|}\) ?
\[
\begin{aligned}
& =\epsilon_{i j k} \frac{\partial}{\partial x_{j}} \frac{J_{k}\left(\vec{r}^{\prime}\right)}{\vec{r}-\vec{r}^{\prime} \mid} \\
& =\epsilon_{i j k} J_{k}\left(\vec{r}^{\prime}\right) \frac{\partial}{\partial x_{j}} \frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|} \\
& =\epsilon_{i k j} J_{k}\left(\vec{r}^{\prime}\right) \frac{x_{j}-x_{j}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}} \\
& =\left[\vec{J}\left(\overrightarrow{r^{\prime}}\right) \times \frac{\vec{r}-\vec{r}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}}\right]_{i}
\end{aligned}
\]
integration and differentiation.

We recover Biot-Savart Law, which is an important consistency check!
\[
\vec{B}=\frac{\mu_{0}}{4 \pi} \int \vec{J}\left(\vec{r}^{\prime}\right) \times \frac{\vec{r}-\vec{r}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}} d^{3} \vec{r}^{\prime}
\]

\section*{The magnetic vector potential : the choice of div A and its consequences}

Our choice of \(A\) cannot affect the final result for \(B\).
But is does effect the solution for BOTH the scalar and the vector potential.
Notice that \(\mathrm{A}, \mathrm{V}\) suffer from "instantaneous change at a distance" problem.
We do not need to care as long as it is a static/steady state solution.
But what if charge and current densities (hence A and V ) are both varying arbitrarily?

The choice of \(A\) and \(V\) : how much freedom is there?
\[
\begin{aligned}
& \nabla \times \vec{E}=-\frac{\partial \vec{B}}{\partial t}=-\frac{\partial}{\partial t} \nabla \times \vec{A} \\
& \nabla \times\left(\vec{E}+\frac{\partial \vec{A}}{\partial t}\right)=0 \\
& \left.\vec{E}+\frac{\partial \vec{A}}{\partial t}=\nabla \text { (some scalar }\right) \\
& \vec{E}=-\nabla V-\frac{\partial \vec{A}}{\partial t} \\
& \vec{E}=-\nabla V-\frac{\partial \vec{A}}{\partial t} \\
& \vec{B}=\nabla \times \vec{A}
\end{aligned}
\]
\[
\begin{aligned}
\vec{A}^{\prime} & =\vec{A}+\nabla \lambda \\
V^{\prime} & =?
\end{aligned}
\]

Suppose \(\nabla V^{\prime}+\frac{\partial \vec{A}^{\prime}}{\partial t}=\nabla V+\frac{\partial \vec{A}}{\partial t}\)
\[
\nabla\left(V-V^{\prime}\right)=\frac{\partial}{\partial t}\left(\vec{A}-\vec{A}^{\prime}\right)=-\frac{\partial}{\partial t} \nabla \lambda
\]

It is possible to set \(\mathrm{V}=0\) and still have an electric field via time varying \(\mathbf{A}\)

V and \(\mathbf{A}\) has to change in such a way that \(E\) and \(B\) remain same.
\(\mathrm{V}, \mathrm{A}\) and \(\mathrm{V}^{\prime}, \mathrm{A}^{\prime}\) will have to be related
\[
\begin{aligned}
\vec{A}^{\prime} & =\vec{A}+\nabla \lambda \\
V^{\prime} & =V-\frac{\partial \lambda}{\partial t}
\end{aligned}
\]
\(\lambda\) is a scalar fn of
\(x, y, z, t\)

Multipole expansion of the magnetic vector potential
\[
\begin{aligned}
& A \vec{r})=\frac{\mu_{0}}{4 \pi} \int \frac{\vec{J}\left(\vec{r}^{\prime}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|} d^{3} \vec{r}^{\prime}=\frac{\mu_{0} I}{4 \pi} \int \frac{d \vec{l}}{\left|\vec{r}-\vec{r}^{\prime}\right|} \\
& \frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|}=\sum_{l=0}^{\infty} \frac{1}{r^{l+1}}\left(r^{\prime}\right)^{l} P_{l}(\cos \theta) \\
& l=0 \quad \frac{\mu_{0} I}{4 \pi} \frac{1}{r} \oint d \vec{l} \quad \text { Always zero: no magnetic monopoles } \\
& l=1 \quad \frac{\mu_{0} I}{4 \pi} \frac{1}{r^{2}} \oint r^{\prime} \cos \theta d \vec{l} \quad \text { Origin independent: magnetic dipole } \\
& l=2 \quad \frac{\mu_{0} I}{4 \pi} \frac{1}{r^{3}} \oint r^{\prime 2} \frac{1}{2}\left(3 \cos ^{2} \theta-1\right) d \vec{l} \quad \text { quadrupole }
\end{aligned}
\]

Multipole expansion of the magnetic vector potential
\[
\begin{aligned}
\oint r^{\prime} \cos \theta d \overrightarrow{r^{\prime}} & =-\frac{1}{2} \hat{r} \times \oint \overrightarrow{r^{\prime}} \times d \overrightarrow{r^{\prime}} \\
\text { Hence } \vec{A}_{\text {dipole }} & =\frac{\mu_{0}}{4 \pi}\left[\frac{1}{2} I \oint \vec{r}^{\prime} \times d \overrightarrow{r^{\prime}}\right] \times \frac{\hat{r}}{r^{2}} \\
& =\frac{\mu_{0}}{4 \pi} \frac{\vec{m} \times \hat{r}}{r^{2}}
\end{aligned}
\]
\(\frac{1}{2} \oint \overrightarrow{r^{\prime}} \times \overrightarrow{d l}=\) area of the loop
dipole moment \(=\) current \(\times\) area

In 3D with volume current density
\(\vec{m}=\frac{1}{2} \int \vec{r} \times \vec{J} \delta \tau\)

\section*{Multipole expansion : an useful identity}

For a localised current distribution (J) with zero divergence at all points, and any two scalar functions \(\mathrm{f}, \mathrm{g}\)
\[
\begin{aligned}
& \int_{v o l} \nabla \cdot(\vec{J} f g) d \tau=\int_{\text {surf }} \vec{J} f g . d \vec{S}=0 \\
& \int_{v o l}[f g(\nabla \cdot \vec{J})+\vec{J} \cdot \nabla f g] d \tau=0 \\
& \int_{v o l}[f \vec{J} \cdot \nabla g+g \vec{J} \cdot \nabla f] d \tau=0
\end{aligned}
\]

Take \(\mathrm{f}=1 \mathrm{~g}=\mathrm{x}, \mathrm{y}, \mathrm{z}\) in turn \& prove

Take \(\mathrm{f}=\mathrm{g}=\mathrm{x}, \mathrm{y}, \mathrm{z}\) in turn \& prove
\[
\int_{\text {vol }} \vec{J} d \tau=0
\]
\[
\int_{v o l} \vec{r} \cdot \vec{J} d \tau=0
\]

Take \(\mathrm{f}=\mathrm{x}, \mathrm{g}=\mathrm{y}\) and other permutaions
\[
\int_{\text {vol }}\left(x J_{y}+y J_{x}\right) d \tau=0
\]

\section*{Multipole expansion of the magnetic vector potential}
\[
\begin{aligned}
A_{i}(\vec{r}) & =\frac{\mu_{0}}{4 \pi} \int \frac{J_{i}\left(\overrightarrow{r^{\prime}}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|} d^{3} \vec{r}^{\prime} \\
& =\frac{\mu_{0}}{4 \pi}\left[\frac{1}{r} \int J_{i}\left(\overrightarrow{r^{\prime}}\right) d^{3} \vec{r}^{\prime}+\frac{1}{r^{3}} \vec{r} \cdot \int \overrightarrow{r^{\prime}} J_{i}\left(\overrightarrow{r^{\prime}}\right) d^{3} \overrightarrow{r^{\prime}}+\cdots\right] \\
& =\frac{\mu_{0}}{4 \pi r^{3}} \vec{r} \cdot \int \overrightarrow{r^{\prime}} J_{i}\left(\vec{r}^{\prime}\right) d^{3} \vec{r}^{\prime}+\cdots \quad \begin{array}{l}
\text { Need to use the identities } \\
\text { derived just before to } \\
\text { obtain the result in the } \\
\text { next step }
\end{array} \\
& =\frac{\mu_{0}}{4 \pi r^{3}}\left[-\frac{1}{2} \vec{r} \times \int \overrightarrow{r^{\prime}} \times \vec{J} d^{3} \vec{r}^{\prime}\right]_{i}+\cdots \\
\vec{A}(\vec{r}) & =\frac{\mu_{0}}{4 \pi} \frac{\vec{m} \times \vec{r}}{r^{3}}+\cdots
\end{aligned}
\]

The magnetic vector potential and field of a perfect dipole
\[
\begin{aligned}
& \vec{A}_{\text {dipole }}=\frac{\mu_{0}}{4 \pi} \frac{m \sin \theta}{r^{2}} \hat{\epsilon}_{\phi} \\
& \vec{B}=\frac{\mu_{0}}{4 \pi} \frac{m}{r^{3}}\left(2 \cos \theta \hat{\epsilon_{r}}+\sin \theta \hat{\epsilon_{\theta}}\right) \\
& \text { Since }\left\{\begin{array}{l}
m \cos \theta=\vec{m} \cdot \hat{\epsilon}_{r} \\
m \sin \theta=\vec{m} \cdot \hat{\epsilon_{\theta}}
\end{array}\right.
\end{aligned}
\]
\(\vec{m}\) lies in the plane defined by \(\left[\epsilon_{r}, \epsilon_{\theta}\right]\)
\(\vec{B}=\frac{\mu_{0}}{4 \pi}\left[\frac{3(\vec{m} \cdot \hat{r}) \hat{r}-\vec{m}}{r^{3}}\right] \quad(r \neq 0)\)
\[
\vec{B}=\frac{\mu_{0}}{4 \pi}\left[\frac{3(\vec{m} \cdot \hat{r}) \hat{r}-\vec{m}}{r^{3}}\right]+\mu_{0} \frac{2}{3} \vec{m} \delta(\vec{r})
\]


If the \(\mathrm{r}=0\) point is to be correctly handled then the delta fn is needed.

Correct solution of the interaction of electron spin and nuclear spin (hyperfine) requires this.

\section*{Forces and torques on current loops and dipoles}

Force on a current distribution will also vanish if all the current loops are closed and the fields are constant
\[
\begin{array}{ll}
\vec{F}=q(\vec{E}+\vec{v} \times \vec{B}) & \text { (single particle) } \\
\delta \vec{F}=(n \delta \tau) q(\vec{E}+\vec{v} \times \vec{B}) & \text { (many particles) } \\
\delta \vec{F}=\vec{J} \times \vec{B} \delta \tau=I \delta \vec{l} \times \vec{B} & \text { (current line, distrib) }
\end{array}
\]

A current carrying wire is electrically neutral because it always has equal number of electrons and positive ions in lattice. An electric field does not create a net force on it.

Magnetic field does, becuase the electrons are moving and the fixed ions in the lattice are not - so the lattice sees no Lorentz force.

Useful facts to remember....
Also \(\nabla . \vec{J}=0\), since \(\frac{\partial \rho}{\partial t}=0\)
\(\oiiint \vec{J} d \tau=0\)
(for a localised charge distribution)... why?

Consider the expression
\(\int_{v o l} \nabla \cdot(x \vec{J}) d \tau=\int[x \nabla \cdot \vec{J}+\vec{J} \cdot \nabla x] d \tau\)
But \(\vec{J} . \nabla x=J_{x}\)
Also \(\vec{J}=0\)
on a large bounding surface
So the result follows

Forces and torques on current loops and dipoles

\(\mathrm{F}=0\)
Torque \(=0\)

What is the force and torque on the square loop?

\[
\vec{F}=\oint I \delta \vec{l} \times \vec{B}
\]

Total \(\mathrm{F}=0\), still, since opposite sides have exactly opposite dl vector
Torque \(=\mathrm{BI} x\) area (in magnitude)
The basis for electric motor winding, pointer type current measuring meters etc.


Only inhomogeneous magnetic field can create a force on a current loop/ dipole.

Torque is possible with uniform fields.

\section*{Forces and torques on current loops and dipoles}

Consider an arbitrary current distribution in a spatially varying field B. Question: What is the force and torque on it?
Assume that the current density is confined to a small volume.
\[
\left.\begin{array}{c}
\vec{F}=\int(\vec{J} \times \vec{B}) d^{3} \vec{r}^{\prime} \text { and } B_{k}(\vec{r})=B_{k}(0)+\vec{r} . \nabla B_{k} \\
F_{i}=\epsilon_{i j k}\left[B_{k}(0) \int J_{j}\left(\vec{r}^{\prime}\right) d^{3} \vec{r}^{\prime}+\int J_{j}\left(\overrightarrow{r^{\prime}}\right) \overrightarrow{r^{\prime}} \cdot \nabla^{\prime} B_{k}(0) d^{3} \vec{r}^{\prime}+\ldots\right] \\
=0
\end{array} \begin{array}{l}
\text { Field inhomogeneity } \\
\text { giving rise to force }
\end{array}\right]
\]

We need to simplify: \(\quad \nabla B_{k}(0) \cdot \int \vec{r}^{\prime} J_{j}\left(\vec{r}^{\prime}\right) d^{3} \vec{r}^{\prime}\)

This will give \(: \quad-\left[\nabla B_{k}(0) \times \frac{1}{2} \int\left(\overrightarrow{r^{\prime}} \times \vec{J}\right) d^{3} \vec{r}^{\prime}\right]_{i}\)
Dipole moment of the current distribution
\[
\vec{F}=(\vec{m} \times \nabla) \times \vec{B}=\nabla(\vec{m} \cdot \vec{B})-\vec{m} \frac{(\nabla \cdot \vec{B})}{=0}
\]

The proof is similar to the one given for the dipole moment calculation before.

The torque will be given by
\[
\begin{aligned}
\vec{\tau} & =\int\left[\overrightarrow{r^{\prime}} \times(\vec{J} \times \vec{B})\right] d^{3} \vec{r}^{\prime} \\
& =\int\left[\vec{J}\left(\overrightarrow{r^{\prime}} \cdot \vec{B}\right)-\vec{B}\left(\overrightarrow{r^{\prime}} . \vec{J}\right)\right] d^{3} \vec{r}^{\prime} \\
& =\vec{B} \cdot \int \overrightarrow{r^{\prime}} \vec{J} d^{3} \overrightarrow{r^{\prime}}+\vec{B} \cdot \int\left(\overrightarrow{r^{\prime}} \cdot \vec{J}\right) d^{3} \overrightarrow{r^{\prime}} \\
& =\vec{m} \times \vec{B} \quad=0
\end{aligned}
\]

Since the zeroth order term does not vanish, we take the value of \(B\) at a fixed point in the distribution and treat it as a constant

Electric dipole
\(\vec{F}=\nabla(\vec{p} \cdot \vec{E})\)
\(\vec{\tau}=\vec{p} \times \vec{E}\)
Magnetic dipole
\(\vec{F}=\nabla(\vec{m} \cdot \vec{B})\)
\(\vec{\tau}=\vec{m} \times \vec{B}\)

The end result is very similar, though the internal mechanism is quite different.

The mechansims by which matter acquires a magnetic "dipole moment" per unit volume is more complex than the way electric polarisation is acquired.

A classical description of this is not really possible.
The magentic moment acquired may be
1. In the direction of the magnetic field but very weak. (paramagnetism)
2. OPPOSITE to the direction of the applied field and also very weak (diamagnetism)

Para \& dia magnetic effects disappear when the applied field is removed.
3. In the direction of the applied field but very strong and remains even after the initial field is removed .(ferromagnetism)
This is characterised by hysteresis effects/loops/

These effects involve the dynamics of orbital electrons of an atom/ free electrons in a metal in a magnetic field, which requires a quantum mechanical description.

We will not focus on "how" the polarisation is acquired.

\section*{Magnetic polarisation and its description}
\[
\vec{A}=\frac{\mu_{0}}{4 \pi} \int_{v o l} \frac{M\left(\vec{r}^{\prime}\right) \times\left(\vec{r}-\vec{r}^{\prime}\right)}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}} d^{3} \vec{r}^{\prime}
\]

Recall that
\[
\nabla_{r^{\prime}} \frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|}=\frac{\vec{r}-\vec{r}^{\prime}}{\left|\vec{r}-\vec{r}^{\prime}\right|^{3}}
\]

Hence
\[
\left.\vec{A}=\frac{\mu_{0}}{4 \pi} \int_{v o l} M \overrightarrow{\left(\vec{r}^{\prime}\right.}\right) \times \nabla_{r^{\prime}} \frac{1}{\left|\vec{r}-\vec{r}^{\prime}\right|} d^{3} \vec{r}^{\prime}
\]
\[
\vec{r}-\vec{r}^{\prime} \quad \nabla X
\]


0

Work out the expansion of \(M \times\) grad using the standard rules. Convert one volume integral of a curl to a surface integral.....
\(\mathbf{M}\) is the magnetic moment per unit volume.

The unit of \(\mathbf{M}\) can be defined in two ways:
[m] = current x area (so Ampere. \(\mathrm{m}^{2}\) )
\([\mathrm{M}]=\mathrm{Am}^{-1}\)
Also [m.B] = energy, hence
[m] = Joule/Tesla
\([\mathrm{M}]=\mathrm{J} /\) Tesla \(/ \mathrm{m}^{-3}\)
But historically an unit emu has been used. emu =1erg/gauss

From which we can define emu/cc or emu/gm

1 erg/gauss \(=10^{-3}\) Ampere. \(\mathrm{m}^{2}\)
Atomic magnetic moments are of the order of a "Bohr magneton"
\[

\]
need to use the relation
\[
\int_{\text {vol }} \nabla \times \vec{A} d \tau=-\int_{\text {surf }} \vec{A} \times d \vec{S}
\]

Contribution from a volume current and a surface current density.


Bar magnel

\(\vec{M}=\) constant
\(\nabla \times \vec{M}=0\)
\(\vec{M} \times \hat{n}=M \hat{\epsilon_{\phi}}\)
mimics the solenoid current equivalnet ampere turns per mt
\[
\begin{aligned}
& \nabla \times \vec{B}=\mu_{0} \vec{J}=\mu_{0}\left(\vec{J}_{f}+\vec{J}_{b}\right)
\end{aligned} \begin{aligned}
& \text { "Free" current put in by wires, } \\
& \text { solenoids etc. }
\end{aligned}
\]

Historically a proportionality between \(\mathbf{M}\) and \(\mathbf{H}\) was emphasized as a material property. This leads to:
\[
\begin{array}{rl|l}
\vec{B} & =\mu_{0}(\vec{H}+\vec{M}) & \\
\vec{M} & =\chi \vec{H} & \\
\vec{B} & =\mu_{0}(1+\chi) \vec{H} & \chi \text { is called susceptibility } \\
\vec{B} & =\mu \vec{H} & \mu \text { is called permeability }
\end{array}
\]

Magnetic polarisation and its description: B H M vectors : Boundary conditions

\[
\begin{aligned}
\overrightarrow{J_{\text {free }}} & =0 & & (\text { at the interface }) \\
\mu_{1} H_{1}^{\perp} & =\mu_{2} H_{2}^{\perp} & & (\text { since } \nabla \cdot \vec{B}=0) \\
H_{1}^{\|} & =H_{2}^{\|} & & (\text {since } \nabla \times \vec{H}=0)
\end{aligned}
\]

The unit of H is same as the unit of M . In SI ampere-turn per meter is generally used. It is dimensionally different from Tesla.

In cgs unit of B and H have same dimensionality. Gauss is used for B Oersted is used for H

Confusion is very common between B and H

Divergence of H is not necessarily zero : An example

Consider a bar magnet with magnetisation M
\[
\begin{aligned}
\text { Air }: \vec{H} & =\frac{\vec{B}}{\mu_{0}} & (M=0) &
\end{aligned} \begin{array}{ll}
\perp \\
& \rightarrow \overrightarrow{\mathrm{air}}=B_{\mathrm{bar}}^{\perp} \\
& \rightarrow
\end{array}
\]
\[
\begin{aligned}
& \nabla \times \vec{H}=0 \\
& \nabla \cdot \vec{H}=-\nabla \cdot \vec{M}=\rho_{m} \\
& \text { is sometimes useful to } \\
& \text { describe an assembly of } \\
& \text { magnets via a potential } \phi \\
& \text { such that } \vec{H}=-\nabla \phi
\end{aligned}
\]

So H can have "sources" and "sinks" like the electric field.

In cases where curl \(\mathbf{H}=0\), it is possible to construct a magnetic scalar potential, whose gradient would give H.

In older texts "magnetic pole density" etc are used. These are the sources and sinks of H , like electric charge is the source and sink of E .

This leads to some confusion about H being the "real" field, which is wrong!!

Physical interpretation of the bound currents

> I

Field of an uniformly magnetised sphere
\(\vec{J}_{b}=\nabla \times M=0\)
\(\vec{\sigma}_{b}=\vec{M} \times \hat{n}=M \sin \theta \epsilon_{\phi}\)
Integrate directly to find \(\vec{A}\) and then \(\vec{B}\)
\(\vec{A}=\frac{\mu_{0} M R^{3}}{3} \frac{\sin \theta}{r^{2}} \hat{\epsilon}_{\phi} \quad(r>R) \quad \vec{B}=\frac{\mu_{0}}{4 \pi}\left(\frac{4 \pi R^{3} M}{3}\right)\left(\frac{2 \cos \theta \hat{\epsilon}_{r}+\sin \theta \hat{\epsilon}_{\theta}}{r^{3}}\right)\)
\(\vec{A}=\frac{\mu_{0} M}{3} r \sin \theta \hat{\epsilon_{\phi}} \quad(r<R) \quad \vec{B}=\frac{2 \mu_{0}}{3} \vec{M}\)
Inside the sphere
\(\vec{H}=\frac{\vec{B}}{\mu_{0}}-\vec{M}=-\frac{\vec{M}}{3}\)
directed opposite to \(\vec{B}\) and \(\vec{M}\)
A somewhat counter-intuitive result!

Field inside is constant
Field outside is that of an equivalent dipole placed at origin.

\section*{Field of a cylindrical bar magnet}


Replace the Magnetisation by an equivalent current: \(\mathrm{M}=\mathrm{NI}\) and calculate the axial field due to all the current loops, using the result for a single loop
\[
B(0,0, z)=\frac{\mu_{0} M}{2}\left[\frac{z+L / 2}{\sqrt{R^{2}+(z+L / 2)^{2}}}-\frac{z-L / 2}{\sqrt{R^{2}+(z-L / 2)^{2}}}\right]
\]

Field just outside \(: B \approx \frac{\mu_{0} M}{2}\)
\[
(z= \pm L / 2)
\]

Field very far away : \(B \approx \frac{\mu_{0}}{2 \pi z^{3}} \cdot\left(M \pi R^{2} L\right) \quad(z \gg L)\)
Calculate H and show that it points opposite to M inside the bar. Typical strong permanent magnets have remnance \(\mu_{0} M \sim 1\) Tesla

\section*{Typical values of susceptibilities}

TABLE 6.1 MAGNETIC SUSCEPTIBILITIES
\begin{tabular}{|lr|}
\hline \multicolumn{1}{|c|}{ Material } & Magnetic Susceptibility \\
\hline Diamagnetic: & \\
Bismuth & \(-16.5 \times 10^{-5}\) \\
Gold & \(-3.0 \times 10^{-5}\) \\
Silver & \(-2.4 \times 10^{-5}\) \\
Copper & \(-0.96 \times 10^{-5}\) \\
Water & \(-0.90 \times 10^{-5}\) \\
Carbon Dioxide & \(-1.2 \times 10^{-8}\) \\
Hydrogen & \(-0.22 \times 10^{-8}\) \\
Paramagnetic: & \(190 \times 10^{-8}\) \\
Oxygen & \(0.85 \times 10^{-5}\) \\
Sodium & \(2.1 \times 10^{-5}\) \\
Aluminum & \(7.8 \times 10^{-5}\) \\
Tungsten & \(48,000 \times 10^{-5}\) \\
Gadolinium &
\end{tabular}

Graphite's suscpetibility can be -6e-4 to -1e-5 depending on orientation in SI units

Source: Handbook of Chemistry and Physics, 67th ed. (Cleveland: CRC Press, Inc., 1986-87.) All figures are for atmospheric pressure and room temperature.

\section*{Stability of levitation: Why does it require diamagnets?}

Since magnetic field repels "diamagnets", it can be made to float in a region of strongly varying (large gradient) magnetic field.


Pieces of graphite floating on a strong magnets. The magnets are typically \(5-10 \mathrm{~mm}\) cubes and would have a remnance of 1-2 Tesla.

The height at which these float are typically \(1-2 \mathrm{~mm}\)


Pictures: Wikipedia
\(\vec{F}=\nabla(\vec{m} \cdot \vec{B})=\rho V g\), where \(\vec{m} \approx V \frac{\chi}{\mu_{0}} \vec{B}\)
Stability requires that \(\nabla \cdot(\nabla \chi \vec{B} \cdot \vec{B})<0\)
But \(\quad \nabla^{2} \vec{B} \cdot \vec{B}=\nabla^{2}\left(B_{x}^{2}+B_{y}^{2}+B_{z}^{2}\right)\)
Show that : \(\sum_{i j} \frac{\partial^{2}}{\partial x_{i}^{2}} B_{j} B_{j}=2 \sum_{j}\left|\nabla B_{j}\right|^{2}\)
Hence \(\quad \nabla^{2} \vec{B} \cdot \vec{B}>0\) always stability requires \(\nabla . \vec{F}<0\) possible only if \(\chi<0\)

Electrodynamics
Something needs to be added to Ampere's Law. Why?
Can we decouple E and B?
Emergence of an wave equation. Why is \(f(x-v t)\) a "wave"?
How does the displacement current term compare with normal current?
Induced emfs: Inductors and generators
Lorentz force Law in potential form (convective derivative)
Energy and momentum of the EM field.
Maxwell's equation in matter
Refractive index
Reflection and transmission of em waves at an interface

Why there must be something more in Ampere's Law.
\(\nabla \times \vec{B}=\mu_{0} \vec{J}\)
If \(\nabla \cdot \vec{J}=0\) then it is fine
Cannot be correct when \(\vec{J}\) is changing
\[
\begin{aligned}
\nabla \cdot \vec{J}+\frac{\partial \rho}{\partial t} & =0 \\
\nabla \cdot \vec{E} & =\frac{\rho}{\epsilon_{0}}
\end{aligned}
\]
always
true


Replace RHS by
\(\vec{J} \rightarrow \vec{J}+\epsilon_{0} \frac{\partial \vec{E}}{\partial t}\)
Full time dependent equation becomes
\(\nabla \times \vec{B}=\mu_{0} \vec{J}+\epsilon_{0} \mu_{0} \frac{\partial \vec{E}}{\partial t}\)
\[
E=\frac{v}{d}=\frac{Q}{C d} \quad Q: \text { chase on the plates }
\]
\(\frac{\partial E}{\partial t} \propto \frac{\partial Q}{\partial t} \propto I(t)\)
The correction term should have something to do with \(\frac{\partial E}{\partial t}\).

Historically the additional term is called "displacement current"

\section*{The full set of Maxwell's equations}
electrostatics
\[
\begin{array}{rlr}
\nabla \cdot \vec{E} & =\frac{\rho}{\epsilon_{0}} & \\
\nabla \times \vec{E} & =-\frac{\partial \vec{B}}{\partial t} & \begin{array}{l}
\text { Faraday } \\
\text { Induction }
\end{array} \\
\nabla \cdot \vec{B} & =0 &
\end{array}
\]
\[
\text { magnetostatics } \quad \nabla \times \vec{B}=\mu_{0} \vec{J}+\epsilon_{0} \mu_{0} \frac{\partial \vec{E}}{\partial t}
\]

Maxwell's displacement current term

These are first order differential equations.
Decoupling them would invariably lead to second order equations.
First do it for free space where there are no charges and currents

Decoupling of Maxwell's equations leads to wave equation in free space
\[
\begin{aligned}
& \nabla \times B=\epsilon_{0} \mu_{0} \frac{\partial E}{\partial t} \text {. } \\
& \nabla \times \nabla \times B=G_{0} \mu, \frac{\partial}{\partial t}(\nabla \times E) \\
& \nabla(\nabla \cdot B)-\nabla^{2} B=\epsilon_{0} \mu_{0} \frac{\partial}{\partial t}\left(-\frac{\partial B}{\partial t}\right) \\
& \theta^{2} B-\epsilon_{0} \mu_{0} \frac{\partial^{2} B}{\partial t^{2}}=0 \\
& \epsilon_{0} \mu_{0}=\frac{1}{c^{2}} \\
& \text { The velocity of light } \\
& \text { emerges "naturally" } \\
& \text { from Maxwell's } \\
& \text { equations } \\
& \nabla \times \nabla \times E=-\frac{\partial}{\partial t}(\nabla \times B) \\
& \theta(\theta \cdot E)-\nabla^{2} E=-\frac{\partial}{\partial t}\left(\epsilon_{0} \mu \cdot \frac{\partial E}{\partial t}\right) \\
& \nabla^{2} E-\epsilon_{0} \mu_{0} \frac{\partial^{2} E}{\partial t^{2}}=0 \\
& \text { What is the generic } \\
& \text { solution of the wave } \\
& \text { equation? }
\end{aligned}
\]

Decoupling of Maxwell's equations leads to wave equation in free space


\section*{Inductors and generators : Self Inductance}

A current carrying loop creates a magnetic field.
Some of the magnetic field creates a flux through the loop.
\(\Phi=L I \quad\) similar to \(\quad V=Q / C\)
How much energy does it take to set this up?
\[
\begin{aligned}
V & =-L \frac{d I}{d t} \quad \text { (Faraday induction) } \\
\delta W & =V \delta Q \quad(\text { work done by the battery) } \\
\frac{d W}{d t} & =V I=\left(\frac{d}{d t} L I\right) I \\
& =\frac{d}{d t}\left(\frac{L I^{2}}{2}\right)
\end{aligned}
\]

Consider a coil of \(N\) turns, With length / and cross sectional area \(A\)
\[
\begin{aligned}
& B=\mu_{0} \frac{N}{l} I \\
& L I=N B A \\
& \frac{(L I) I}{2}=\frac{N B A}{2} \frac{B l}{\mu_{0} N} \\
& W=\frac{B^{2}}{2 \mu_{0}} A l \quad \\
& \quad \text { Energy per unit volume } \\
& \text { in magnetic field }
\end{aligned}
\]

In a circuit capacitor stores energy in its electric field. The inductor stores energy in its magnetic field.
More generally...
Energy in the electromagnetic field \(\int_{\text {vol }} d \tau\left[\frac{\epsilon_{0} E^{2}}{2}+\frac{B^{2}}{2 \mu_{0}}\right]\)

\section*{Inductors and generators : Mutual Inductance}

Flux through loop 2 due to loop 1
\[
\begin{aligned}
\Phi_{2} & =\int \vec{B}_{1} \cdot d \vec{S}_{2}=\int\left(\nabla \times \vec{A}_{1}\right) \cdot d \vec{S}_{2} \\
& =\int \vec{A}_{1} \cdot \overrightarrow{d l_{2}}=\int \frac{I_{1} \overrightarrow{d l_{1}}}{\left|\vec{r}_{1}-\vec{r}_{2}\right|} \cdot \int \overrightarrow{d l}_{2} \\
M I_{1}= & I_{1} \oint \oint \frac{\overrightarrow{d l}_{1} \cdot \overrightarrow{d l}_{2}}{\left|\vec{r}_{1}-\vec{r}_{2}\right|} \quad \begin{array}{l}
\text { Mutual Inductance } \\
\text { (notice the symmetry) }
\end{array}
\end{aligned}
\]


Question:
\[
\begin{aligned}
\Phi_{1} & =L I_{1}+M I_{2} \\
& \text { OR } \\
\Phi_{1}= & L I_{1}-M I_{2}
\end{aligned}
\]

The flux due to coil 2 may increase or decrease the flux through coil 1 That depends on the geometry and assumed direction of the instantaneous current

The dot convention : If both currents are entering the dot then fluxes will add.
Equivalently: If current enters the dotted terminal then the polarity of the voltage at the other dotted terminal will be positive.

\section*{Inductors and generators : Mutual Inductance : Energy stored in the full system}

How much is the energy stored in the system L1 + L2 ?
\(W=L_{1} I_{1}^{2}+L_{2} I_{2}^{2}+M I_{1} I_{2} \quad\) When all currents either enter or leave the dot
\(W=L_{1} I_{1}^{2}+L_{2} I_{2}^{2}-M I_{1} I_{2} \quad\) When one current enters and the other leaves the dot

This follows from the relation between dot and the addition/subtraction of flux
Can be generalised to an arbitrary number of coupled linear inductors

\section*{Inductors and generators : The simple generator}

http://misswise.weebly.com/motors-and-generators.html
Either the magnet OR coil may rotate

If the magnet rotates, the change in flux is obvious
\(\frac{d}{d t} \Phi=\frac{d}{d t} N \cdot B \cdot A \cdot \cos \omega t\)
\(V=-N . B \cdot A \cdot \omega \sin \omega t\)

N turns in the coil Area of each turn : A angular freq of rotation : \(\omega\)


Notice that the voltage WILL alternate or have ripples even if the contacts to the terminals are flipped every half cycle

Inductors and generators : The simple generator : coil rotates, magnet stationary
\[
\oint_{\text {coil }} \vec{E} \cdot \overrightarrow{d l}=\oint_{\text {coil }} \vec{v} \times \vec{B} \cdot \overrightarrow{d r} \quad \text { (Lorenz Force) } \quad \begin{aligned}
& \text { Noties the } \mathbf{B} \text { vector } \\
& \text { is fixed in this case }
\end{aligned}
\]

For rigid rotation : \(\vec{v}=\vec{\omega} \times \vec{r} \quad\) so \(\quad(\vec{\omega} \times \vec{r}) \times \vec{B}=\vec{r}(\vec{B} \cdot \vec{\omega})-\vec{\omega}(\vec{B} \cdot \vec{r})\)
\(\vec{B} \cdot \vec{\omega}=0 \quad \& \quad \vec{B} \cdot \vec{r}=B r \sin \omega t \quad(\theta=\omega t\) is the angle with the normal to the coil \()\)
The integral again gives \(V=-B . A \cdot \omega \cdot \sin \omega t\)

Example: Conducting rod rolling on a track.....(Lorenz force or flux change??)


So "flux changes due to magnet motion" and "circuit moves" cases gives the same result. [See refs for interesting comments]

The usual viewpoint is: Faraday induction is correct in the frame where circuit is at rest. But this is not always easy to apply...becuase the circuit may deform, expand or stretch (motion, but not centre of mass motion.)

EITHER apply Faraday's law, treating the derivative carefully OR apply Lorentz Force argument

References:
Feynman Lect vol2, chap16 J D Jackson Chap 6

Inductors and generators : The NOT SO SIMPLE homopolar generator


Force on the moving charge : \(\vec{F}=q \vec{v} \times \vec{B}\) In travelling from center to edge of the disk
\[
\begin{aligned}
\int \frac{\vec{F}}{q} \cdot \overrightarrow{d l} & =\int_{0}^{r} \vec{v} \times \vec{B} \cdot \overrightarrow{d l} \\
& =B \omega \frac{r^{2}}{2}
\end{aligned}
\]
\[
\begin{aligned}
& \vec{B}=B \hat{\epsilon}_{z} \\
& \vec{v}=\omega r \hat{\epsilon}_{\theta}
\end{aligned}
\]

Faraday disk/ homopolar/unipolar generator.
Question: Where is the change of flux?
Why does it generate a voltage at all?
Observation: Rotating the disk generates a voltage. But rotating the magnet does NOT.

How is the "circuit" being completed? What is the path of the electron?

uniform magnetic field \(\boldsymbol{B}\) acting into the page

How does someone sitting in the disk explain the voltage generated?

Faraday induction vs Lorenz force. Are they equivalent?
circuit is moving through \(\vec{B}(x, y, z)\)
which has no \(t\) dependence Loop/circuit may not be rigid... parts may have diff speeds

Since \(\vec{B}\) has no \(t\) dependence
\(\oiint \vec{B} \cdot \vec{d} a=0 \quad\) always \(/\) any closed surface \(R+S+R^{\prime}\)
\(\Phi(t+\delta t)-\Phi(t)=\oiint_{R^{\prime}} \vec{B} \cdot \overrightarrow{d a}+\oiint_{R} \vec{B} \cdot \overrightarrow{d a}\)
\[
\delta \Phi=-\oiint \vec{B} \cdot d \vec{a}
\]
\[
=-\oint_{R} \vec{B} \cdot \vec{d} l \times \vec{v} \delta t
\]
\[
-\frac{d \Phi}{d t}=\oint_{R} \vec{v} \times \vec{B} \cdot \overrightarrow{d l}
\]


When R is considered part of the closed surface the direction of the outward normal reverses. So the sign also revereses. Important!

But this will not work if B has explicit time dependence because the surfaces \(R\) and \(R^{\prime}\) are not traced out at equal time, we cannot set integral of B.da to zero over the surface.

But Faraday's law is true irrespective of whether B is time dependent or not.

The two laws are NOT equivalent, but of course consistent with each other!
'Voltage'' in a non-conservative field situation
Non conservative fields (non zero curl) can give rise to puzzling situations. An example: flux changes in the central region and induced a current in the \(R+R+R\) loop.


Two voltmeters are connected across the same two points.
Q: Should they not measure the same "voltage"?
A: NO, not in this case.
But....if the top lead of \(A\) is disconnected and laid along the red dotted line, then both will measure the same. (figure out why..)

What do "voltmeters" measure? Faraday's law in a multiply connected region.
R H Romer, American JI of Physics 50, 1089 (1982)

\section*{Energy and Momentum of particle + EM field system}

Conservative field \(\rightarrow\) KE + PE (scalar potential) conserved.
EM fields are in general not conservative, so what is conserved?
Expectation: KE of particles + "something" will be conserved.
\[
\delta W_{M}=\int_{\text {allvol }} \rho(\vec{E}+\vec{v} \times \vec{B}) \cdot \vec{v} \delta t d \tau
\]
\[
\frac{d W_{M}}{d t}=\int \vec{E} \cdot \vec{j} d \tau
\]
\[
=\frac{1}{\mu_{0}} \int(\vec{E} \cdot \nabla \times \vec{B}) d \tau-\frac{\partial}{\partial t} \int \frac{\epsilon_{0} E^{2}}{2} d \tau
\]
\[
=-\frac{1}{\mu_{0}} \int \nabla \cdot(\vec{E} \times \vec{B}) d \tau+\frac{1}{\mu_{0}} \int \vec{B} \cdot(\nabla \times \vec{E}) d \tau-\frac{\partial}{\partial t} \int \frac{\epsilon_{0} E^{2}}{2} d \tau
\]
\[
=-\frac{1}{\mu_{0}} \int \nabla \cdot(\vec{E} \times \vec{B}) d \tau-\frac{\partial}{\partial t} \int\left(\frac{\epsilon_{0} E^{2}}{2}+\frac{B^{2}}{2 \mu_{0}}\right) d \tau
\]

Hence \(\frac{d}{d t}\left[W_{M}+\int_{\text {vol }}\left(\frac{\epsilon_{0} E^{2}}{2}+\frac{B^{2}}{2 \mu_{0}}\right) d \tau\right]=-\frac{1}{\mu_{0}} \int_{\text {surf }} \nabla \cdot(\vec{E} \times \vec{B}) d \tau\)
compare with \(\nabla \cdot \vec{j}+\frac{\partial \rho}{\partial t}=0\) Hence \(\frac{d Q_{\text {in }}}{d t}=-\int_{\text {surf }} \vec{j} \cdot d \vec{a}\)
\[
\frac{d}{d t}\left[W_{M}+\int_{\text {vol }}\left(\frac{\epsilon_{0} E^{2}}{2}+\frac{B^{2}}{2 \mu_{0}}\right) d \tau\right]=-\int_{\text {surf }} \frac{\left.\frac{1}{\mu_{0}}(\vec{E} \times \vec{B}) \cdot d \vec{a}\right]}{l}
\]

Energy of particles + field

Work done on the charges=
Force x displacement (integrated over all vol)
Use \(\nabla \times \vec{B}=\mu_{0} \vec{j}+\epsilon_{0} \mu_{0} \frac{\partial \vec{E}}{\partial t}\) to replace \(\vec{j}\)
\(\nabla \cdot(\vec{E} \times \vec{B})=\vec{B} \cdot \nabla \times \vec{E}-\vec{E} . \nabla \times \vec{B}\) The energy flux

\section*{Energy and Momentum of particle + EM field system}

We find that the EM field contains energy and we can identify the energy flux/flow/current term as well.

Natural question: Can we do the same for momentum of the particles? This is more invloved, becuase momentum is a vector and forming the continuity equation for a vector would require a "tensor".

Apart from that the reasoning is very similar...
\[
\begin{aligned}
\frac{d}{d t} \sum_{\text {all }} \vec{p}_{i} & =\vec{F}=\int_{\text {all vol }} \rho(\vec{E}+\vec{v} \times \vec{B}) d \tau \\
& =\int\left[\left(\epsilon_{0} \nabla \cdot \vec{E}\right) \vec{E}+\left(\frac{\nabla \times \vec{B}}{\mu_{0}}-\epsilon_{0} \frac{\partial \vec{E}}{\partial t}\right) \times \vec{B}\right] d \tau
\end{aligned}
\]

Since \(:(\nabla \times \vec{B}) \times \vec{B}=(\vec{B}, \nabla) \vec{B}-\nabla \frac{B^{2}}{2}\)
\[
\text { And : } \begin{aligned}
\left(\frac{\partial \vec{E}}{\partial t}\right) \times \vec{B} & =\frac{\partial}{\partial t}(\vec{E} \times \vec{B})+\vec{E} \times(\nabla \times \vec{E}) \\
& =\frac{\partial}{\partial t}(\vec{E} \times \vec{B})-\left[(\vec{E} . \nabla) \vec{E}-\nabla \frac{E^{2}}{2}\right]
\end{aligned}
\]
we have used Faraday's law
\(\nabla \times \vec{E}=-\frac{\partial \vec{B}}{\partial t}\)
and a similar expansion again

\section*{Energy and Momentum of particle + EM field system}

RHS becomes :
\(\epsilon_{0}\left[(\nabla . \vec{E}) \vec{E}+(\vec{E} . \nabla) \vec{E}-\nabla \frac{E^{2}}{2}\right]+\frac{1}{\mu_{0}}\left[(\nabla . \vec{B}) \vec{B}+(\vec{B} . \nabla) \vec{B}-\nabla \frac{B^{2}}{2}\right]-\frac{1}{c^{2}} \frac{\partial}{\partial t} \frac{(\vec{E} \times \vec{B})}{\mu_{0}}\)

The integrand is now remarkably symmetric in E and B although the initial expression was not. The extra term we have added is div B which is always zero.

\section*{\(\mathbf{S}=\mathrm{ExB}\)}
emerges again
\[
\begin{aligned}
\frac{d}{d t}\left[\sum_{\text {particles }} \vec{p}_{i}+\frac{1}{c^{2}} \int \vec{S} d \tau\right]=\int & {\left[\epsilon_{0}\left\{(\nabla \cdot \vec{E}) \vec{E}+(\vec{E} \cdot \nabla) \vec{E}-\nabla \frac{E^{2}}{2}\right\}+\right.} \\
& \left.\frac{1}{\mu_{0}}\left\{(\nabla \cdot \vec{B}) \vec{B}+(\vec{B} \cdot \nabla) \vec{B}-\nabla \frac{B^{2}}{2}\right\}\right] d \tau
\end{aligned}
\]

Question : Is RHS the divergence of something? Then the form of the continuity equation will emerge again.

But the RHS is already a vector, so it can only be the divergence of tensor (if at all)

\section*{Energy and Momentum of particle + EM field system}
\[
\begin{aligned}
& {\left[(\nabla \cdot \vec{E}) \vec{E}+(\vec{E} \cdot \nabla) \vec{E}-\nabla \frac{E^{2}}{2}\right]_{i} } \\
= & \frac{\partial E_{j}}{\partial x_{j}} E_{i}+E_{j} \frac{\partial E_{i}}{\partial x_{j}}-\frac{1}{2} \frac{\partial E^{2}}{\partial x_{i}} \\
= & \frac{\partial}{\partial x_{j}}\left(E_{i} E_{j}-\delta_{i j} \frac{E^{2}}{2}\right)
\end{aligned}
\]

Repeated index \(j\) is summed over, there is no summation over i

Hence the entire RHS integrand is a divergence of the following quantity \(T_{i j}=\epsilon_{0}\left(E_{i} E_{j}-\delta_{i j} \frac{E^{2}}{2}\right)+\frac{1}{\mu_{0}}\left(B_{i} B_{j}-\delta_{i j} \frac{B^{2}}{2}\right)\)

Formally called the Electromagnetic (Maxwell) stress tensor
\(\frac{d}{d t}\left[\sum_{\text {particles }} \vec{p}_{i}+\frac{1}{c^{2}} \int \vec{S} d \tau\right]=-\int_{\text {vol }} \nabla \cdot(-\underline{\underline{T}}) d \tau=-\int_{\text {surf }}(-\underline{\underline{T}}) \cdot d \vec{a}\) compare with \(\quad \frac{d}{d t} Q_{\text {inside }}=-\int_{\text {vol }} \nabla \cdot \vec{j} d \tau \quad=-\int_{\text {surf }} \vec{j} \cdot d \vec{a}\)

Q: Why would you call it a stress tensor?

Energy and Momentum of particle + EM field system
\[
\frac{d}{d t} \sum_{\text {particles }} \vec{p}_{i}=\left[-\frac{1}{c^{2}} \frac{d}{d t} \int \vec{S} d \tau+\int_{\text {surf }} \underline{T} \cdot d \vec{a}\right]
\]

If we take the volume to include all the particles (or a solid object) then the RHS tells us the total force on that volume.

If there is no \(t\) dependence then the integral of \(T\) gives the force. In a "mechanical" or "fluid" situation, this is exactly what the stress tensor would have given us.

This formulation can also be used to analyse cases where a focussed beam of light is used to hold up a particle...."optical tweezer"..

b

DNA
2000000000000000000
Extension

Bead

Laser beam

Physics world: Optical tweezers: where physics meets biology : Nov 13, 2008

Electromagnetic Waves in free space: What is a plane EM wave?
\[
\begin{aligned}
& \vec{E}=\vec{E}_{0} e^{i(\vec{k} \cdot \vec{r}-\omega t)} \\
& \vec{B}=\vec{B}_{0} e^{i(\vec{k} \cdot \vec{r}-\omega t)}
\end{aligned}
\]

We start by assuming that Ko and Bo have no spatial dependence, they do NOT depend on \(x, y, z\). All spatial dependence comes from the exponential.

Of course En and Bo can have \(x, y, z\) components, but they are all constants. These must satisfy Maxwell's equations.
\[
\begin{aligned}
& \begin{array}{lll}
\nabla \cdot \vec{E} & =0 \\
\nabla \cdot \vec{B} & =0
\end{array} \longrightarrow \\
& \nabla \times \vec{E}=-\frac{\partial \vec{B}}{\partial t} \\
& \nabla \times \vec{B}=\frac{1}{c^{2}} \frac{\partial \vec{E}}{\partial t} \\
& \text { Hence } \vec{k} \cdot \vec{E}_{0}=0 \\
& \text { \& } \quad \vec{k} \cdot \vec{B}_{0}=0
\end{aligned}
\]
\[
\begin{aligned}
& \nabla \cdot \vec{E}_{0} e^{i(k \cdot r-\omega t)} \\
= & \frac{\partial}{\partial x} E_{0 x} e^{i(k \cdot r-\omega t)}+\frac{\partial}{\partial y} E_{0 y} e^{i(k \cdot r-\omega t)} \ldots \\
= & i\left(E_{0 x} k_{x}+E_{0 y} k_{y}+E_{0 z} k_{z}\right) e^{i(k \cdot r-\omega t)} \\
= & i \vec{k} \cdot \vec{E}_{0} e^{i(k \cdot r-\omega t)}
\end{aligned}
\]

We would not have got this "transverse" condition without the assumption of Ko and Bo being constant....in waveguides the condition does NOT hold.

Electromagnetic Waves in free space : What is a plane EM wave?
The third equation gives:
\(\vec{k} \times \vec{E}_{0}=\omega \vec{B}_{0}\)
Hence
\[
\begin{aligned}
\vec{E}_{0} \times\left(\vec{k} \times \vec{E}_{0}\right) & =\omega \vec{E}_{0} \times \vec{B}_{0} \\
\vec{k}\left(\vec{E}_{0} \cdot \vec{E}_{0}\right)-\vec{E}_{0}\left(\vec{k} \cdot \vec{E}_{0}\right) & =\omega \vec{E}_{0} \times \vec{B}_{0} \\
\vec{k} & =\omega \frac{\vec{E}_{0} \times \vec{B}_{0}}{E_{0}^{2}} \\
\left|B_{0}\right| & =\frac{\left|E_{0}\right|}{c}
\end{aligned}
\]

The wave propagates in the direction of \(E \times B\).

The relative magnitudes:
For reasonably strong \(\mathrm{E}=1000 \mathrm{~V} / \mathrm{m}\) \(B \sim 3\) microTesla very weak .

That's why we mostly talk about coupling with the electric field of light.

\section*{Electromagnetic Waves in free space : Wavefronts and their shapes}

Wavefront of plane waves

\section*{Plane normal to \(\mathbf{k}\)}
k.r \(=\) length of the red line \(x\) magnitude of \(k\) as long as the tip of \(\mathbf{r}\) lies in the plane.
Surfaces of constant k.r at a certain time \(t\) are called wavefronts. For plane waves the wavefronts are planes.
For sphereical waves these would be spherical surfaces.

Simple spherical wavefront described by
\[
\begin{aligned}
& V(r, t)=\frac{A}{r} e^{i(k r-\omega t)} \\
& \text { It is NOT } \vec{k} \cdot \vec{r}-\omega t
\end{aligned}
\]

Any wave coming from a source (like light from a point) is in reality spherical. But at large distances it is approximated by a plane wave very well.

This is similar to neglecting the earth's curvature over a small region....

\section*{Electromagnetic Waves in free space :Energy, momentum density and Intensity}
\[
\vec{E}(r, t)=\vec{E}_{0} \cos (\vec{k} \cdot \vec{r}-\omega t)
\]

Hence \(\left\langle E^{2}\right\rangle=\frac{1}{T} \int_{0}^{T} E_{0}{ }^{2} \cos ^{2}(\vec{k} \cdot \vec{r}-\omega t) d t\)
\(\begin{aligned} & =\frac{E_{0}{ }^{2}}{2} \\ \text { Energy } \quad U & =\left(\frac{\epsilon_{0}\left\langle E^{2}\right\rangle}{2}+\frac{\left\langle B^{2}\right\rangle}{2 \mu_{0}}\right)=\frac{\epsilon_{0} E_{0}{ }^{2}}{2}\end{aligned}\)
Momentum \(\vec{p}=\frac{\vec{S}}{c^{2}}=\frac{1}{\mu_{0} c^{2}}\langle\vec{E} \times \vec{B}\rangle\)
\[
|p|=\frac{U}{c}
\]

Using the earlier results
\[
\begin{aligned}
|B| & =\frac{|E|}{c} \\
\epsilon_{0} \mu_{0} & =\frac{1}{c^{2}}
\end{aligned}
\]

Although the \(B\) field is much weaker, E and B components make equal contributions to the field energy.


Intensity : Energy passing through per unit area per unit time.
Intentsity \(\quad I=\frac{A(c \delta t) U}{A \delta t}=U c\)

All the energy in the volume will pass through the cross section in time dt

Maxwell's equation in 'linear" matter : what happens to the wave equation?
We consider an insulator first, so there are no free charges in the material
\[
\begin{aligned}
& \vec{D}=\epsilon \vec{E} \\
& \vec{B}=\mu \vec{H}
\end{aligned}
\]
\[
\begin{aligned}
& \nabla \cdot \vec{D}=0 \\
& \nabla \cdot \vec{B}=0
\end{aligned}
\]

But now both magnetisation and electric polarisation can simultaneously change. So the "bound" current will result from change in \(\mathbf{M}\) as well as \(\mathbf{P}\).
\[
\sigma_{b}=\vec{P} \cdot \hat{n} \quad: \quad \text { Then consider } \vec{P} \rightarrow \vec{P}+\overrightarrow{\delta P}
\]

This change causes some amount of charge to flow in/out
\[
\begin{aligned}
& \delta Q=\delta(\vec{P} \cdot \hat{n}) \delta a \\
& \vec{J}_{p} \cdot \overrightarrow{\delta a}=\frac{\delta Q}{\delta t}=\frac{\partial \vec{P}}{\partial t} \cdot \overrightarrow{\delta a} \\
& \nabla \times \vec{B}=\mu_{0} \vec{J}_{\text {total }}+\epsilon_{0} \mu_{0} \frac{\partial \vec{E}}{\partial t} \\
& \nabla \times\left[\mu_{0}(\vec{H}+\vec{M})\right]=\mu_{0}\left[\vec{J}_{f}+\nabla \times \vec{M}+\frac{\partial \vec{P}}{\partial t}\right]+\mu_{0} \frac{\partial}{\partial t}[\vec{D}-\vec{P}] \\
& \nabla \times \vec{H} \quad=\quad \vec{J}_{f}+\frac{\partial \vec{D}}{\partial t} \\
& \text { Show that this } \\
& \text { interpretation is } \\
& \text { consistent with the } \\
& \text { continuity } \\
& \text { equation }
\end{aligned}
\]

Maxwell's equation in 'linear' matter : what happens to the wave equation?
\[
\begin{aligned}
\nabla \cdot \vec{D} & =0 \\
\nabla \cdot \vec{B} & =0 \\
\nabla \times \vec{H} & =\vec{J}_{f}+\frac{\partial \vec{D}}{\partial t} \\
\nabla \times \vec{E} & =-\frac{\partial \vec{B}}{\partial t}
\end{aligned}
\]
\[
\begin{aligned}
\vec{B} & =\mu_{0}(\vec{H}+\vec{M}) \\
\epsilon_{0} \vec{E} & =(\vec{D}-\vec{P})
\end{aligned}
\]
\[
\vec{D}=\epsilon \vec{E}
\]
\[
\vec{B}=\mu \vec{H}
\]

With \(\vec{J}_{f}=0 \quad\) we will get \(\quad \nabla \times \vec{B}=\mu \epsilon \frac{\partial \vec{E}}{\partial t}\)
The wave will propagate with speed \(\quad v^{2}=\frac{1}{\mu \epsilon}\)
Refractive index of the medium \(\quad n=\frac{c}{v}=\sqrt{\frac{\mu \epsilon}{\mu_{0} \epsilon_{0}}}\)

\section*{Maxwell's equation in 'linear' matter : The boundary conditions}

Consider a boundary between two media 1 and 2
Since \(\operatorname{div} \mathrm{D}=0\), the normal component of D must be continuous. div \(B=0\), always (so normal component of \(B\) is continuous)

Since curl H has no singularities ... the tangential component of H is continuous curl \(E\) has no singularities ... the tangnetial component of \(E\) is continuous
\[
\begin{aligned}
D_{1}^{\perp} & =D_{2}^{\perp} \quad \text { Hence } \epsilon_{1} E_{1}^{\perp}=\epsilon_{2} E_{2}^{\perp} \\
B_{1}^{\perp} & =B_{2}^{\perp} \\
H_{1}^{\|} & =H_{2}^{\|} \quad \text { Hence } \frac{B_{1}^{\|}}{\mu_{1}}=\frac{B_{2}^{\|}}{\mu_{2}} \\
E_{1}^{\|} & =E_{2}^{\|}
\end{aligned}
\]

These boundary conditions govern the reflection and transmission of electromagnetic waves at an interface and hence the laws of reflection and refraction (optics)

\section*{Electromagnetic waves at an interface : reflection and transmission}

Normal incidence
\[
x=0
\]

The incident wave propagating to the right
\[
\begin{aligned}
\vec{E}_{I} & =E_{0 \mathrm{I}} e^{i\left(k_{1} x-\omega t\right)} \hat{y} \\
\vec{B}_{I} & =\frac{1}{v_{1}} E_{0 \mathrm{I}} e^{i\left(k_{1} x-\omega t\right)} \hat{z}
\end{aligned}
\]

The reflected wave propagating to the left
\[
\begin{aligned}
\vec{E}_{R} & =E_{0 \mathrm{R}} e^{i\left(-k_{1} x-\omega t\right)} \hat{y} \\
\vec{B}_{R} & =-\frac{1}{v_{1}} E_{0 \mathrm{R}} e^{i\left(-k_{1} x-\omega t\right)} \hat{z}
\end{aligned}
\]

The transmitted wave propagating to the right
\[
\begin{aligned}
\vec{E}_{R} & =E_{0 \mathrm{R}} e^{i\left(k_{2} x-\omega t\right)} \hat{y} \\
\vec{B}_{R} & =\frac{1}{v_{2}} E_{0 \mathrm{R}} e^{i\left(k_{2} x-\omega t\right)} \hat{z}
\end{aligned}
\]

Tangential E Tangential H are continuous
\[
\begin{aligned}
E_{0 \mathrm{I}}+E_{0 \mathrm{R}} & =E_{0 \mathrm{~T}} \\
\frac{1}{\mu_{1}}\left(\frac{E_{0 \mathrm{I}}}{v_{1}}-\frac{E_{0 \mathrm{R}}}{v_{1}}\right) & =\frac{1}{\mu_{2}} \frac{E_{0 \mathrm{~T}}}{v_{2}} \\
\text { define } \quad \beta & =\frac{\mu_{1} v_{1}}{\mu_{2} v_{2}}
\end{aligned}
\]

Need to solve for the ratios only....
\[
\begin{aligned}
& \frac{E_{0 \mathrm{R}}}{E_{0 \mathrm{I}}}=\frac{1-\beta}{1+\beta}=\left|\frac{n_{1}-n_{2}}{n_{1}+n_{2}}\right| \\
& \frac{E_{0 \mathrm{~T}}}{E_{0 \mathrm{I}}}=\frac{2}{1+\beta}=\left(\frac{2 n_{1}}{n_{1}+n_{2}}\right)
\end{aligned}
\]

\section*{Electromagnetic waves at an interface : reflection and refraction}

An useful result with three "phasor" s:
\(A e^{i a x}+B e^{i b x}=C e^{i c x} \quad \forall x\)
Then \(a=b=c\)
set \(x=0 \quad\) : this gives \(\quad A+B=C\)
This condition determines the length of the phasors, which must be satisfied at all times

Now draw the three phasors when \(x \neq 0\)
Two sides of a traingle are together greater than the third side

The equality can only hold if
\(A, B, C\) are along the same ray..
The phase angle also must be same


Electromagnetic waves at an interface : reflection and refraction
\[
\begin{aligned}
& \vec{E}_{R}=\vec{E}_{0 \mathrm{R}} \exp \left[i\left(\vec{k}_{R} \cdot \vec{r}-\omega t\right)\right] \quad \text { Oblique incidence at an interface (general case) } \\
& \vec{B}_{R}=\frac{\hat{k}_{R} \times \vec{E}_{0 \mathrm{R}}}{v_{1}} \exp \left[i\left(\vec{k}_{R} \cdot \vec{r}-\omega t\right)\right] \\
& \vec{E}_{T}=\quad \vec{E}_{\text {оТ }} \exp \left[i\left(\overrightarrow{k_{T}} \cdot \vec{r}-\omega t\right)\right] \\
& \vec{B}_{T}=\frac{\hat{k}_{T} \times \vec{E}_{0 \mathrm{~T}}}{v_{2}} \exp \left[i\left(\vec{k}_{T} \cdot \vec{r}-\omega t\right)\right] \\
& \vec{E}_{I}=\quad \vec{E}_{0 I} \exp \left[i\left(\vec{k}_{I} \cdot \vec{r}-\omega t\right)\right] \\
& \text { at the } \mathrm{x}=0 \text { plane } \\
& \text { sideways } \\
& \vec{B}_{I}=\frac{\hat{k}_{I} \times \vec{E}_{0 I}}{v_{1}} \exp \left[i\left(\vec{k}_{I} \cdot \vec{r}-\omega t\right)\right] \\
& \text { Notice how unit vectors have } \\
& \text { been used to fix the relative } \\
& \text { directions }
\end{aligned}
\]
\(\omega=|\vec{k}| v:\) Hence \(k_{I} v_{1}=k_{R} v_{1}=k_{T} v_{2} \quad\) Use the result \(\vec{k}_{I} \cdot \vec{r}=\vec{k}_{R} \cdot \vec{r}=\vec{k}_{T} \cdot \vec{r} \quad\) must hold \(\forall r\) on the \(x=0\) plane
\[
\begin{aligned}
& k_{I}=k_{R}=\frac{v_{2}}{v_{1}} k_{T} \quad \text { in magnitude } \\
& \left.\begin{array}{l}
\left(k_{I}\right)_{y} y+\left(k_{I}\right)_{z} z=\left(k_{R}\right)_{y} y+\left(k_{R}\right)_{z} z \\
\left(k_{I}\right)_{y} y+\left(k_{I}\right)_{z} z=\left(k_{T}\right)_{y} y+\left(k_{T}\right)_{z} z
\end{array}\right\} \text { holds } \forall y, z
\end{aligned}
\]


This means all the coefficients ( \(y, z\) components) must be equal
Form the triple product of \(k_{l}, k_{R}, k_{T}\) : this must vanish since two row/columns are identical.
The three vectors are co-planer [Law of reflection and refraction] Let this be the \(x\) - \(y\) plane.

Since \(\left|\boldsymbol{k}_{\boldsymbol{l}}\right|=\left|\boldsymbol{k}_{\boldsymbol{R}}\right|\) and y components are equal, the other ( x ) component is exactly reversed. No other possibility can satisfy all these conditions.

Equality of the \(y\)-components require
\(k_{I} \sin \theta_{I}=k_{R} \sin \theta_{R}=k_{T} \sin \theta_{T}\)
\[
\theta_{I}=\theta_{R}
\]
\[
\frac{\sin \theta_{I}}{\sin \theta_{T}}=\frac{v_{1}}{v_{2}}=\frac{n_{2}}{n_{1}}
\]```

